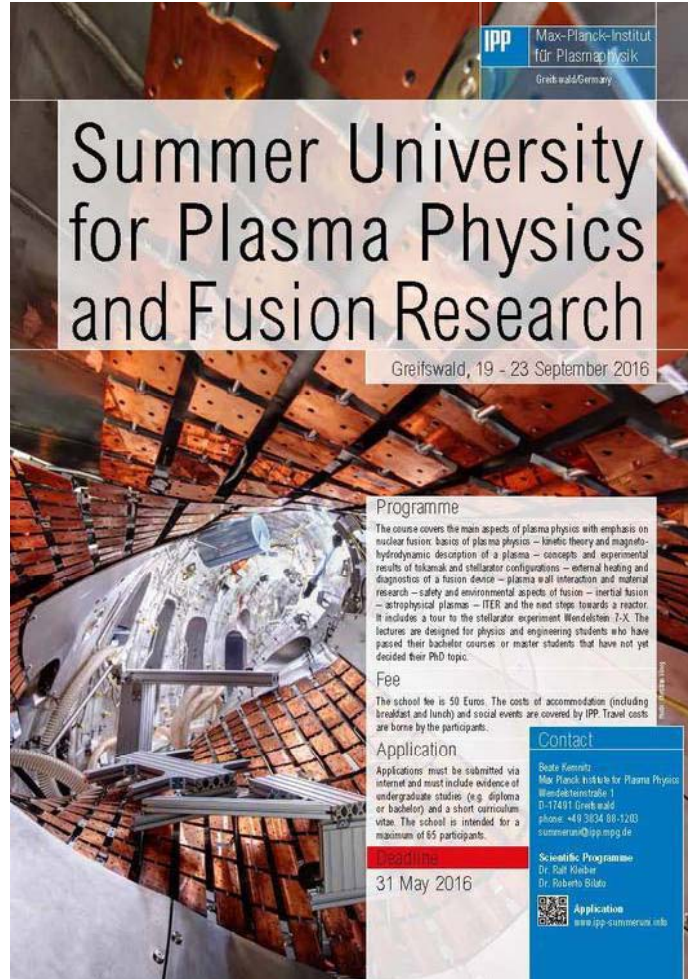




# Welcome to Greifswald



**ipp** Max-Planck-Institut  
für Plasmaphysik  
Greifswald/Germany

## Summer University for Plasma Physics and Fusion Research

Greifswald, 19 - 23 September 2016

### Programme

The course covers the main aspects of plasma physics with emphasis on nuclear fusion: basics of plasma physics – kinetic theory and magneto-hydrodynamic description of a plasma – concepts and experimental results of tokamak and stellarator configurations – external heating and diagnostics of a fusion device – plasma wall interaction and material research – safety and environmental aspects of fusion – inertial fusion – astrophysical plasmas – ITER and the next steps towards a reactor. It includes a tour to the stellarator experiment Wendelstein 7-X. The lectures are designed for physics and engineering students who have passed their bachelor courses or master students that have not yet decided their PhD topic.

### Fee

The school fee is 50 Euros. The costs of accommodation (including breakfast and lunch) and social events are covered by IPP. Travel costs are borne by the participants.

### Application

Applications must be submitted via internet and must include evidence of undergraduate studies (e.g. diploma or bachelor) and a short curriculum vitae. The school is intended for a maximum of 65 participants.

### Deadline

31 May 2016


### Contact

Silke Kemnitz  
Max-Planck-Institut für Plasmaphysik  
Wendelsteinstraße 1  
D-17491 Greifswald  
phone: +49 3834 86-1203  
summeruni@ipp.mpg.de

### Scientific Programme

Dr. Ralf Kniebe  
Dr. Roberto Sato

### Application

 [www.ipp-summeruni.org](http://www.ipp-summeruni.org)



# Plasma physics in Greifswald



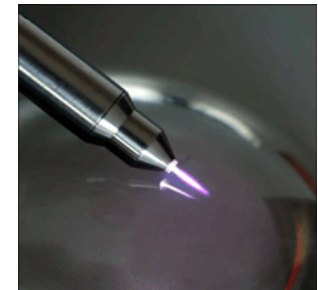
Max-Planck-Institut für Plasmaphysik

- fusion research



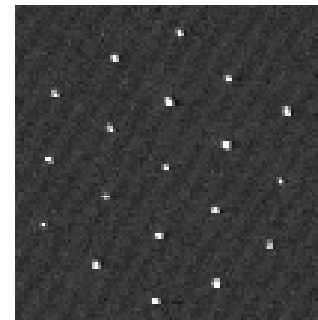
Leibniz-Institut für Plasmaforschung

- plasma technology



Institut für Physik Universität Greifswald

- academic research





# Basic plasma physics and the motion of charged particles in magnetic fields

**Andreas Dinklage**  
*Max-Planck-Institut für Plasmaphysik (IPP)*

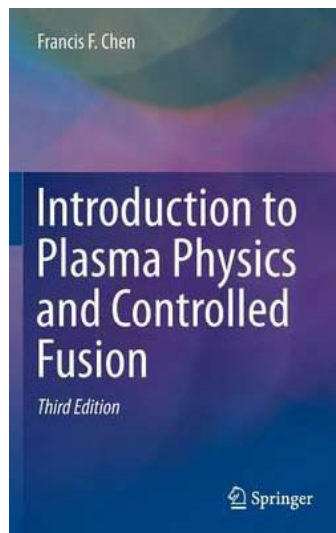




# Outline



- ① What is  $\pi\lambda\sigma\mu\alpha$ ? The (A) world of plasmas.
- ② Overview: what makes a plasma a plasma?
- ③ How to tame plasmas for fusion: motion in magnetic fields



Introduction to Plasma Physics and Controlled Fusion  
by Francis F. Chen  
Springer (3<sup>rd</sup> edition, 2016)





# Outline



- ① What is *πλάσμα*? The (A) world of plasmas.
- ② Overview: what makes a plasma a plasma?
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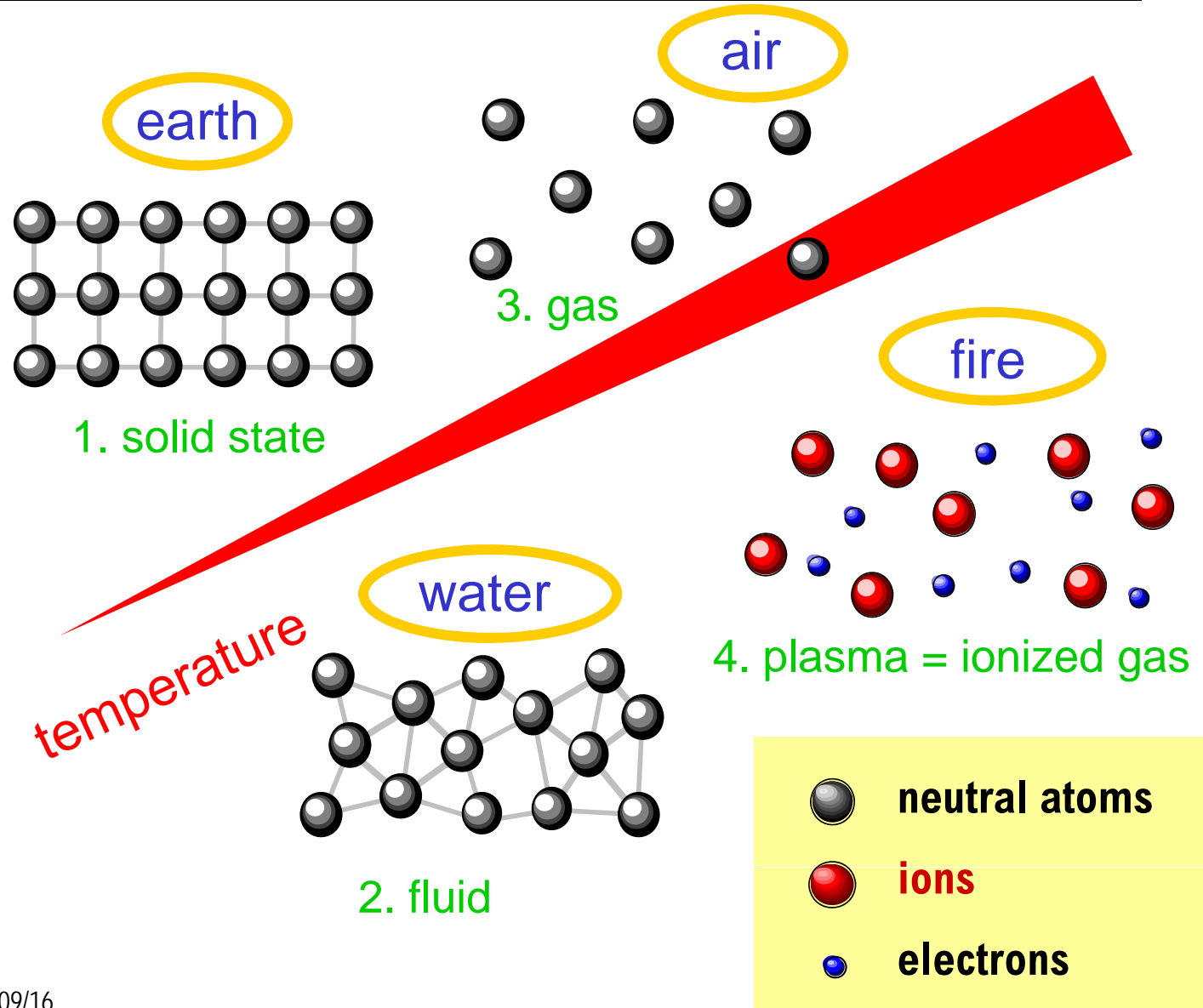


# A world of plasmas



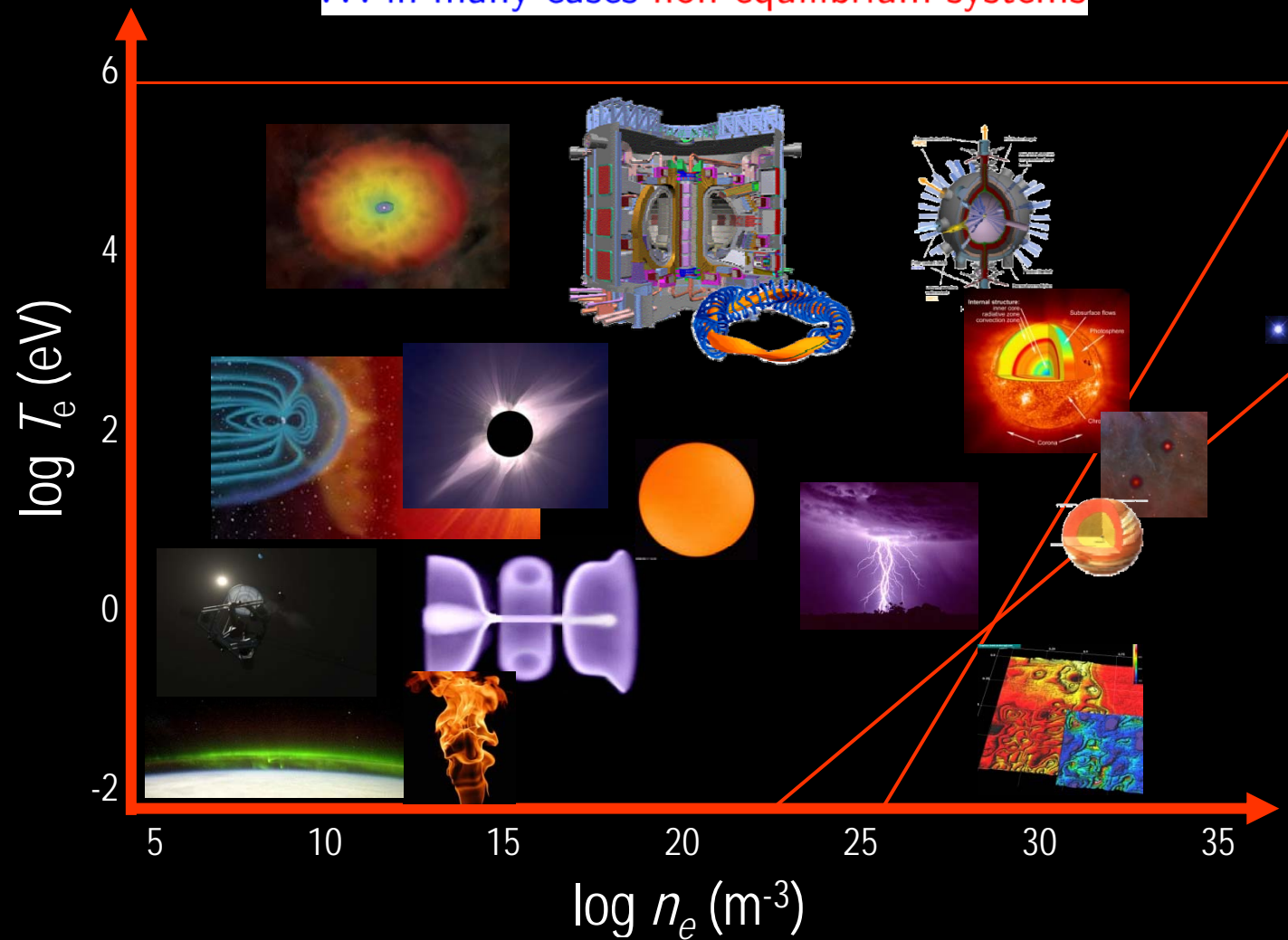
Empedokles  
494-432 B.C

**World  
consists of  
four elements**



# The world of plasmas

... in many cases non-equilibrium systems



$1\text{eV} = e/k_B \text{ K} \approx 11.600 \text{ K}$



# Outline



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# Outline

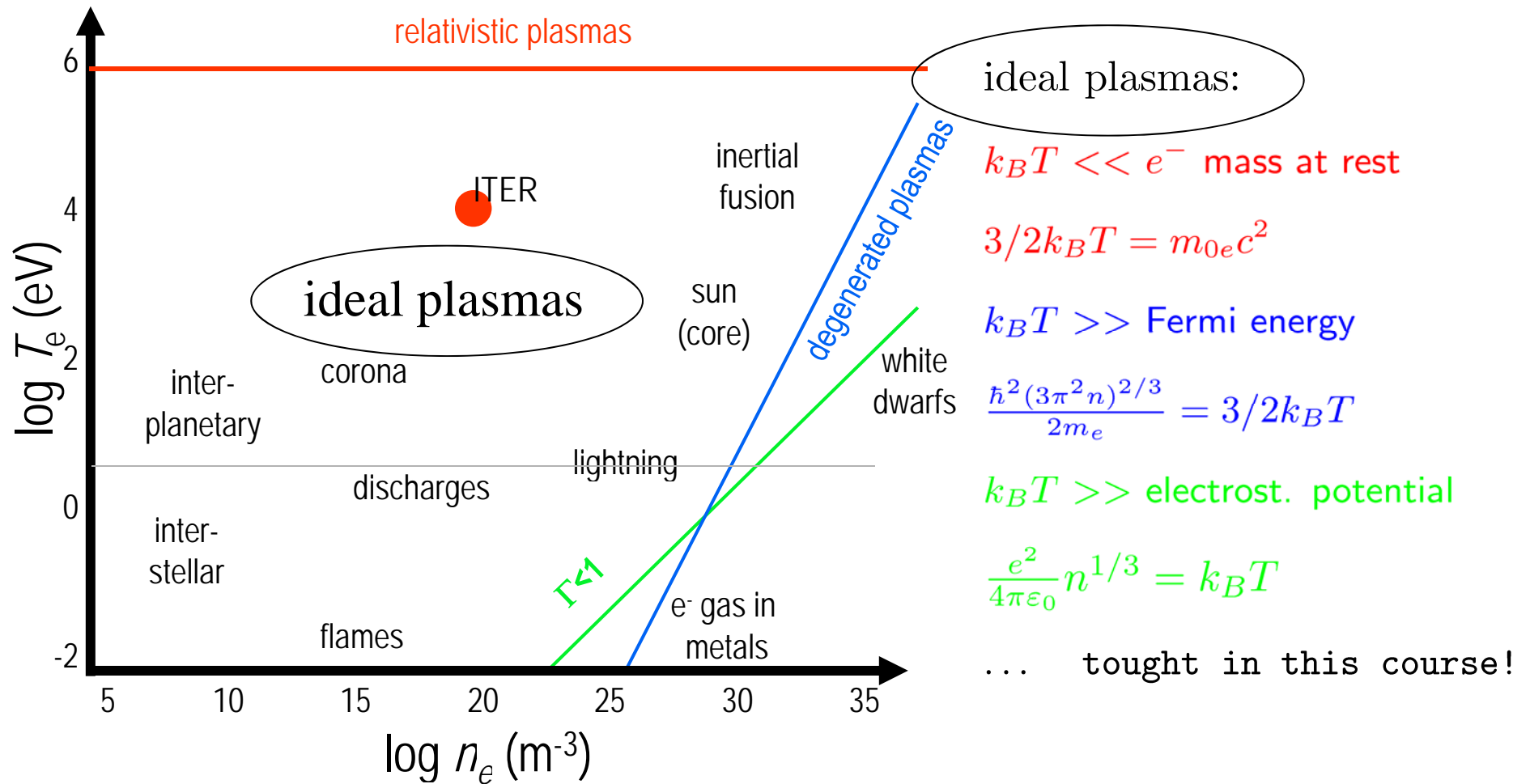


- ① What is  $\pi\lambda\alpha\sigma\mu\alpha$ ? The (A) world of plasmas.
- ② Overview: what makes a plasma a plasma?
  - ① limits: relativity, quantum effects, ionization ...
  - ② time scales - plasma frequency
  - ③ lengths - Debye shielding
  - ④ Coulomb collisions and transport quantities
- ③ How to tame plasmas for fusion: motion in magnetic fields





# Plasma map

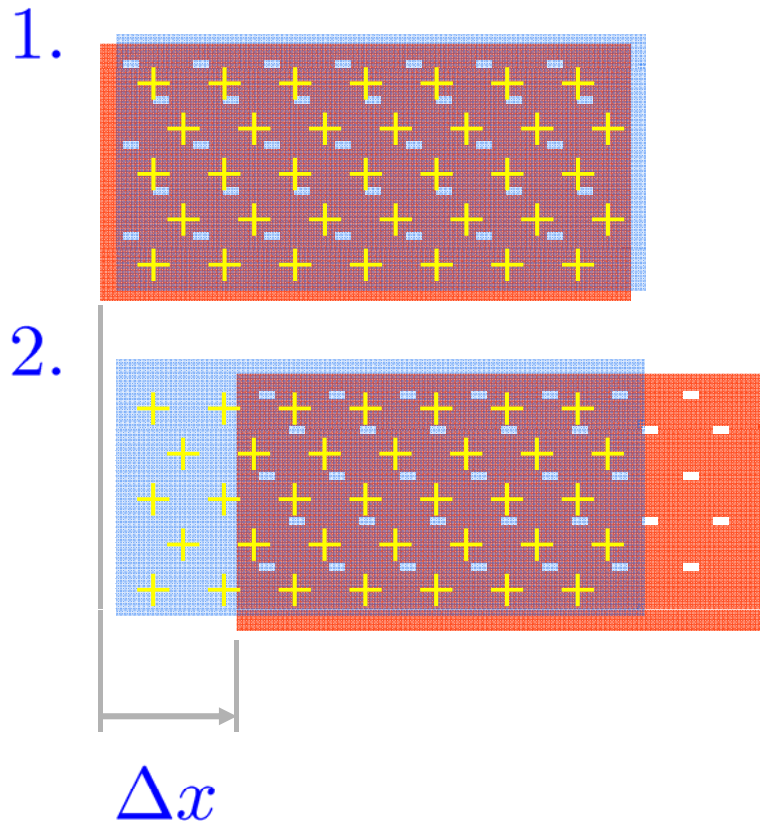




# Plasma frequency



## *Gedankenexperiment*



displacement of charges

⇒ electric field

⇒ repelling force

⇒ oscillation

$$E = \frac{\sigma}{\epsilon_0} = \frac{e n_e \Delta x}{\epsilon_0}$$

figures:

$$\Delta x = 1 \text{ mm}$$

$$n_e = 1 \times 10^{20} \text{ m}^{-3}$$

$$\Rightarrow E \approx 1.8 \times 10^9 \text{ Vm}^{-1}$$

quasi-neutrality



# Plasma frequency



ions at rest, small perturbation:

+ equation of motion

$$m_e \frac{\partial v_1}{\partial t} = -eE_1$$

+ continuity equation

$$\frac{\partial n_1}{\partial t} + n_e \frac{\partial v_1}{\partial x} = 0$$

+ Poisson's equation

$$\epsilon_0 \frac{\partial E_1}{\partial x} = -e n_1$$

+ harmonic Ansatz  $(\omega, k)$ :

$$-i m_e \omega v_1 = -i \frac{n_e e^2}{\epsilon_0 \omega} v_1$$

•  $v_1$  ( $T$ ) cancels out,  $n_e$  not!

•  $\partial\omega/\partial k = 0$ : Langmuir oscillation

$$\omega_p = \left( \frac{n_e e^2}{\epsilon_0 m_e} \right)^{1/2}$$

$$f_p (\text{Hz}) \approx 9 \sqrt{n (\text{m}^{-3})}$$

figures:

$$n_e = 1 \times 10^{20} \text{ m}^{-3}$$

$$\Rightarrow f_p \approx 90 \text{ GHz}$$



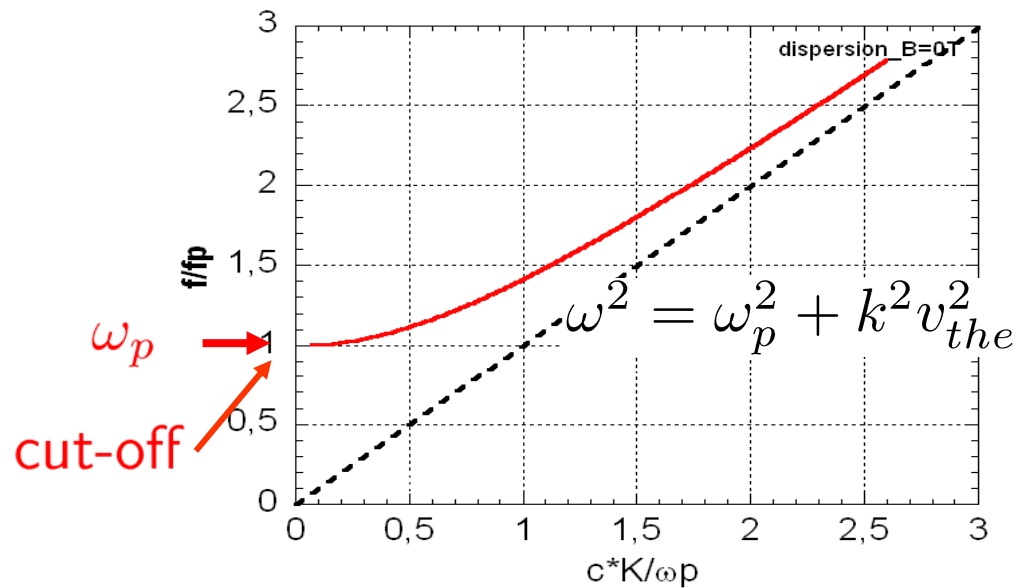
# Plasma frequency: Apps.



wave propagation and reflection

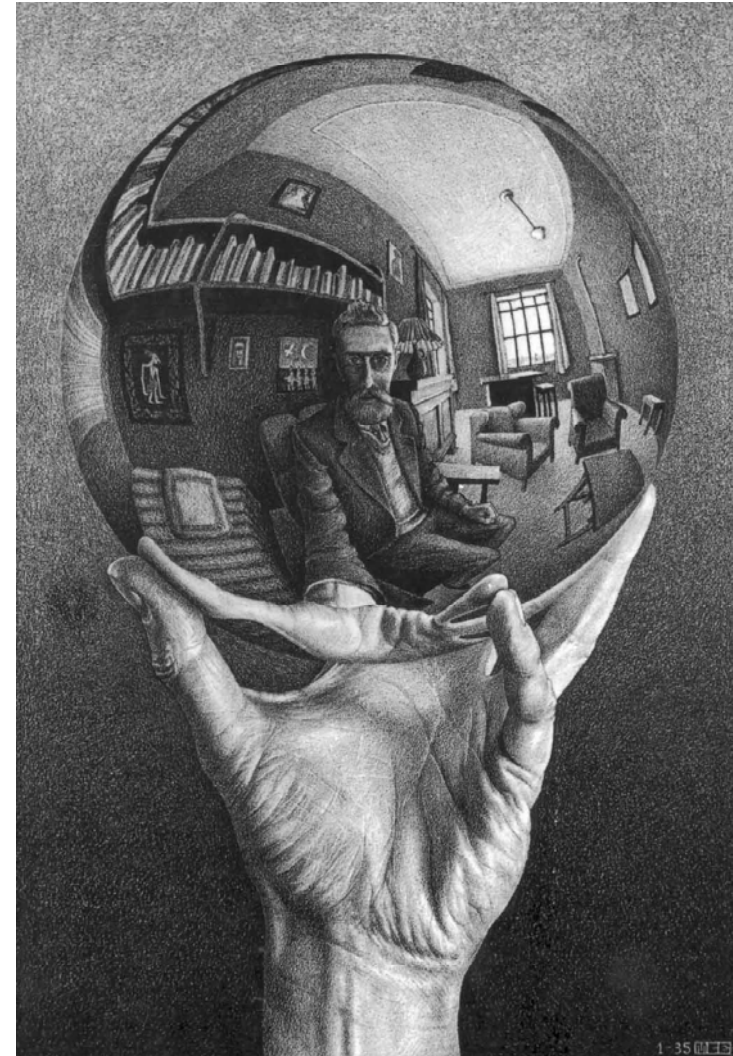
dispersion  $\omega = \omega(k)$

include thermal motion of electrons



in metals, e.g.  $n_e^{Ag} \approx 6 \times 10^{28} \text{ m}^{-3}$

$\lambda_{prop} < 135 \text{ nm}$ , i.e. reflection above



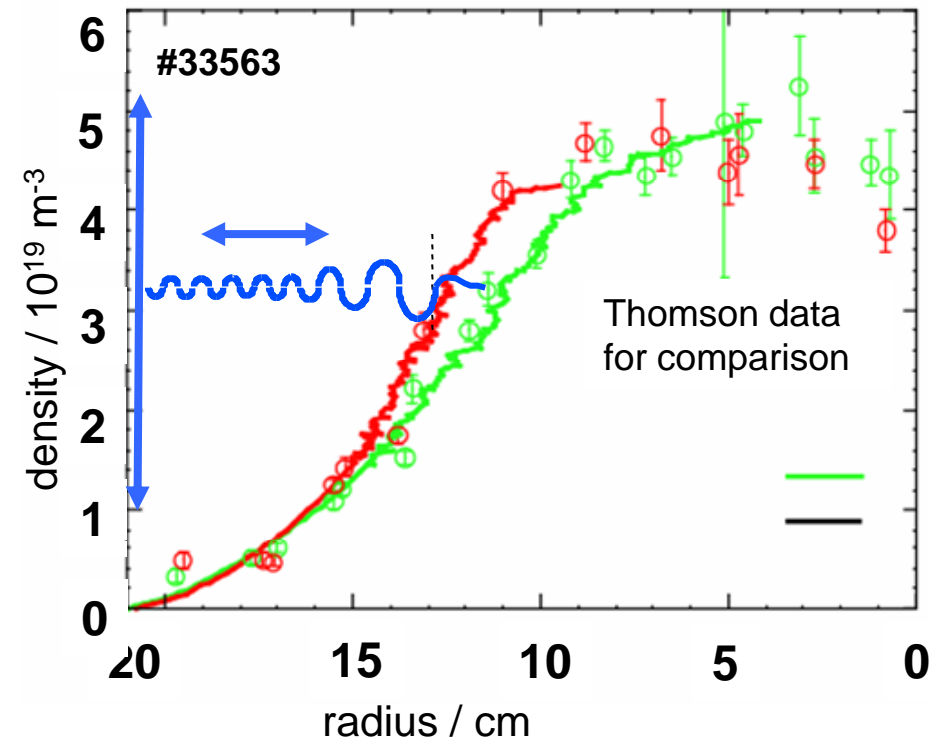
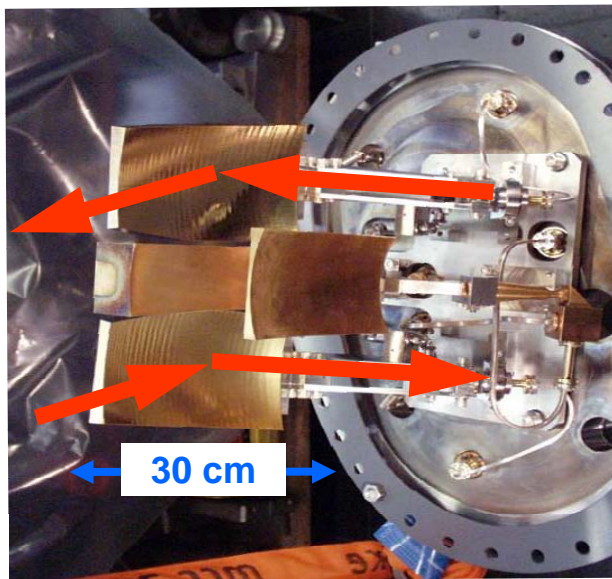


# Plasma frequency: Apps.



Reflectometry:

Density profiles: radar-like (time-of-flight) detection of cut-off layer





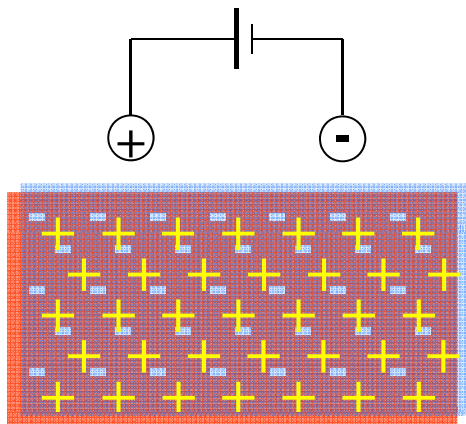


# Debye shielding

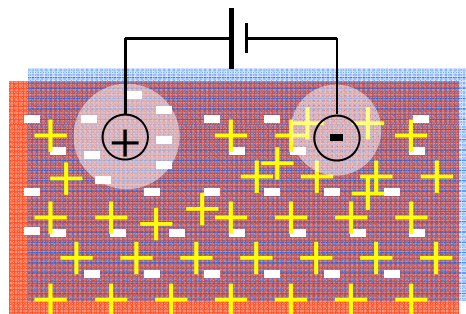


## Gedankenexperiment

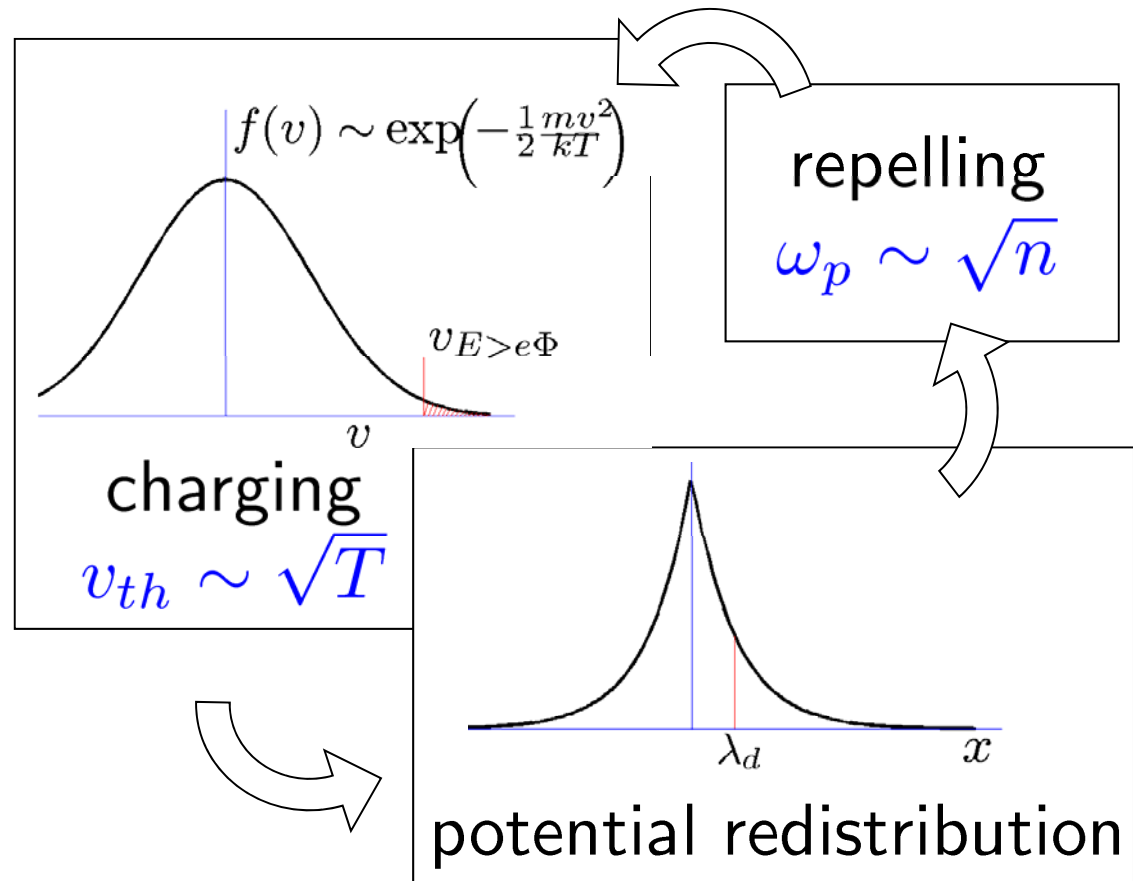
1.



2.



perturbation in a plasma



⇒ shielding



# Debye length



shielding:

heavy ions ( $M/m_e = \infty$ ):

+ Maxwellian:

$$f(v) \sim \exp\left(-\frac{1/2mv^2 + q\Phi}{k_B T_e}\right)$$

→ Boltzmann distribution

$$n_e = n_\infty \exp\left(\frac{q\Phi}{k_B T_e}\right)$$

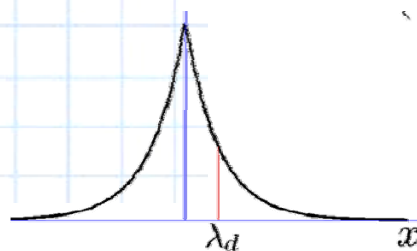
in Poisson's equation

$$\varepsilon_0 \frac{d^2 \Phi}{dx^2} = -e(n_i - n_e)$$

gives Debye-Hückel potential

$$\Phi = \Phi_0 \exp\left(-\frac{|x|}{\lambda_d}\right)$$

• sheath develops



$$\lambda_d = \left(\frac{\varepsilon_0 k_B T_e}{n e^2}\right)^{1/2}$$

$$\lambda_d(m) \approx 7430 \sqrt{\frac{T(eV)}{n(m^{-3})}} \quad (1)$$

figures:

$$n_e = 1 \times 10^{20} \text{ m}^{-3}$$

$$T_e = 10 \text{ keV}$$

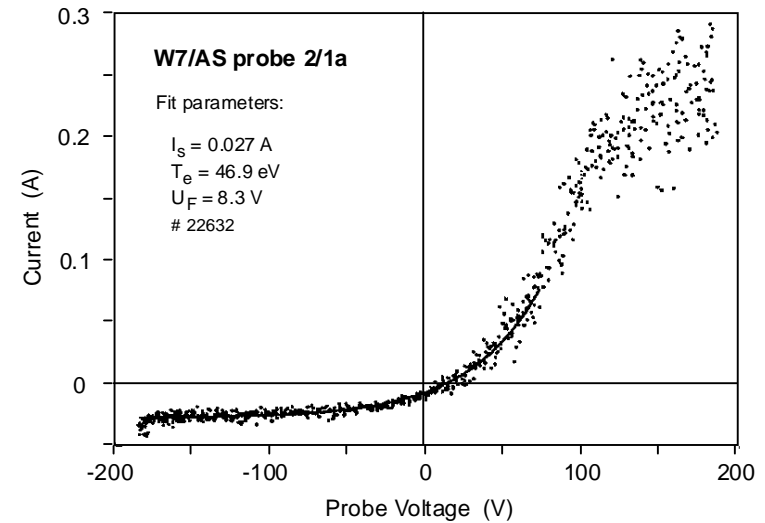
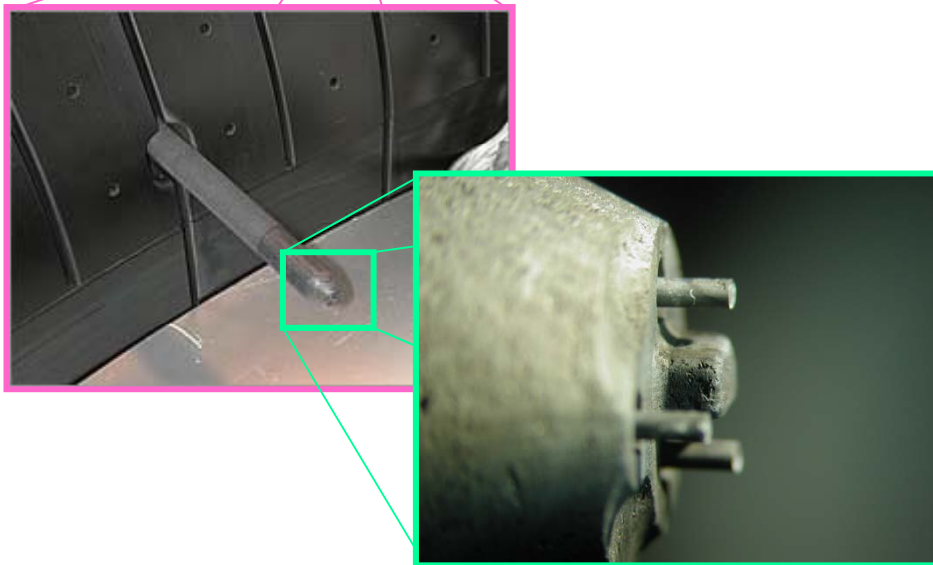
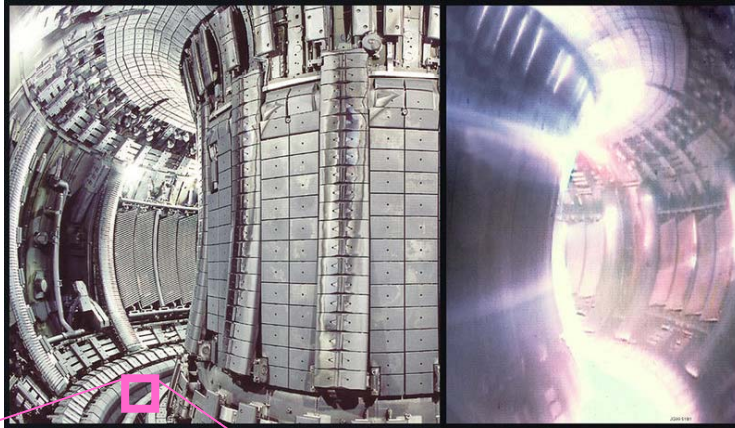
$$\Rightarrow \lambda_d \approx 75 \mu\text{m}$$

$$\begin{aligned} (1) \quad 1\text{eV} &= e/k_B \text{ K} \approx 11.000 \text{ K} \\ e \cdot T(\text{eV}) &= (k_B T) \text{ (J)} \end{aligned}$$

Dinklage: *Basic Plasma Physics* (16/47)



# Debye length: Apps.



Langmuir probe  
current-voltage characteristics:

- electron density  $n_e$
- electron temperature  $T_e$

in 'cold' plasmas

© JET-EFDA



# Outlook: turbulence in magnetized plasmas



## Probe arrays



**Turbulence** affects transport, stability, ...  
the efficiency of a power plant

**Turbulence** is topical, basic science

**Turbulence** affected by

- external magnetic fields
- internal electric fields
- plasma boundaries

*Example: propagation of turbulent structures*

lifetime

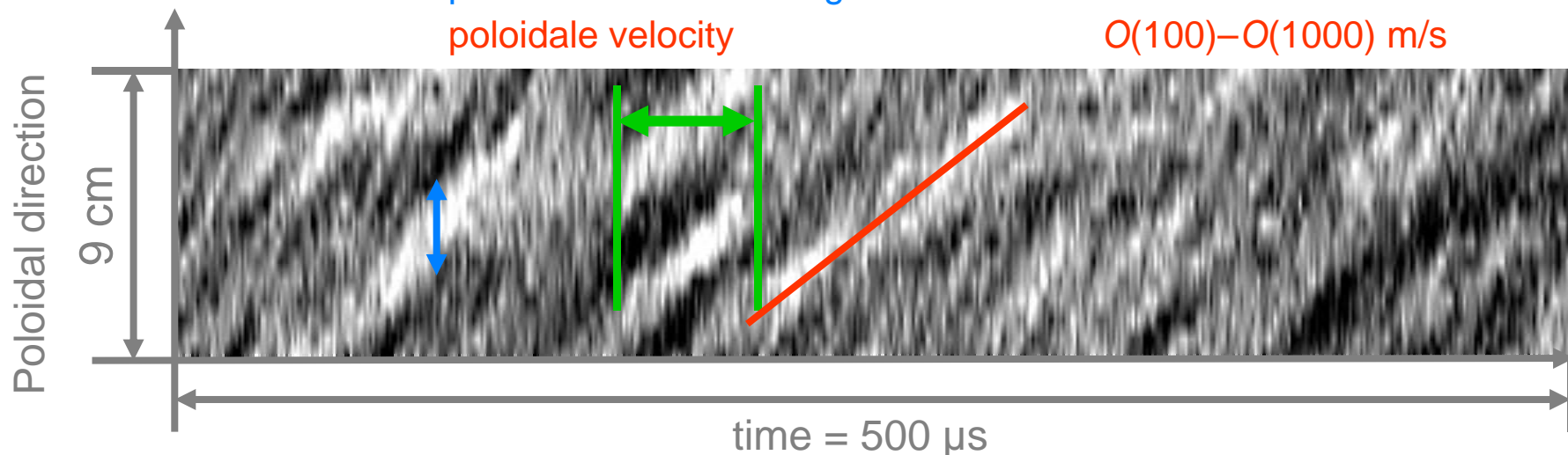
about  $10 \mu\text{s}$

poloidal correlation length

1–5 cm

poloidal velocity

$O(100)$ – $O(1000)$  m/s





Plasmas are quasi-neutral many-particle systems.

Collective effects of charged particles are conveyed by electromagnetic fields.

Criteria:

1.  $\lambda_d \ll L, n/\nabla n, T/\nabla T, \dots$
2.  $N_d \gg \gg 1$  ( $N_d = 4\pi/3 n \lambda_d^3$ ) ( $\Gamma_c = kT/E_{pot} > 1$ )
3.  $\omega_{Plasma} \tau_{collisions} > 1$





# Collisions & elementary processes



## 3. $\omega_{Plasma} \tau_{collisions} > 1$

ionization 3-body recombination	$e^- + A^{+Z}$	$\rightarrow$ $\leftarrow$	$e^- + A^{+Z+1} + e^-$
collisional excitation coll. de-excitation	$e^- + A^{+Z}$	$\rightarrow$ $\leftarrow$	$e^- + (A^{Z+})^*$
photo ionization radiative recombination	$h\nu + A^{+Z}$	$\rightarrow$ $\leftarrow$	$(A^{+Z+1})^* + e^-$
absorption spontaneous emission	$h\nu + A^{+Z}$	$\rightarrow$ $\leftarrow$	$(A^{Z+})^*$
absorption Bremsstrahlung	$h\nu + e^- + A^{+Z}$	$\rightarrow$ $\leftarrow$	$A^{Z+} + (e^-)^*$
Coulomb collisions	$e^- + A^{+Z}$	$\rightleftharpoons$	$(A^{Z+})' + (e^-)'$
fusion e.g.	$D + T$	$\rightleftharpoons$	$n + \alpha + 17.6 \text{ MeV}$



# Collisions in hot fusion plasmas



Coulomb collisions	$e^- + A^{+Z}$	$\rightleftharpoons$	$(A^{Z+})' + (e^-)'$
fusion e.g.	$D + T$	$\rightleftharpoons$	$n + \alpha + 17.6 \text{ MeV}$

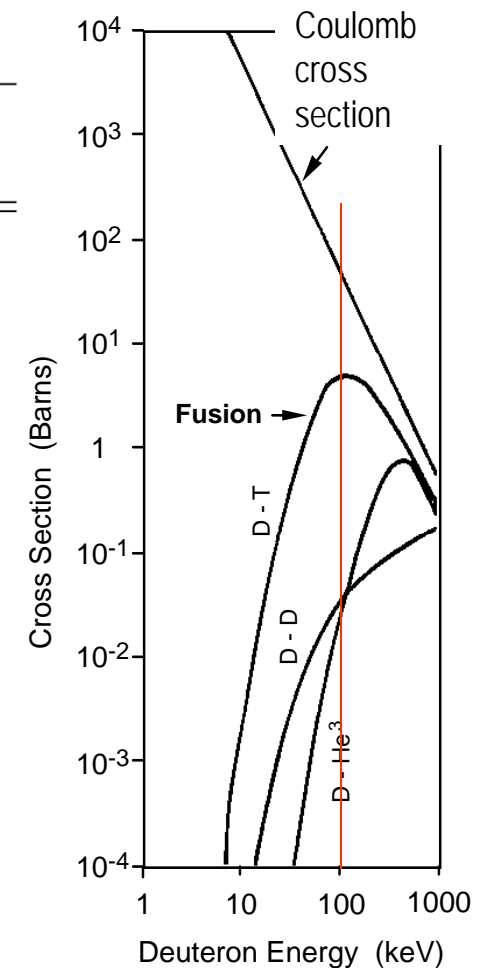
Effect of collisions? Reaction rate.

single collision - cross section:  $\sigma \text{ (m)}^2$

ensemble - rate coefficient :  $\langle \sigma v \rangle \text{ (m}^3\text{s}^{-1}\text{)}$

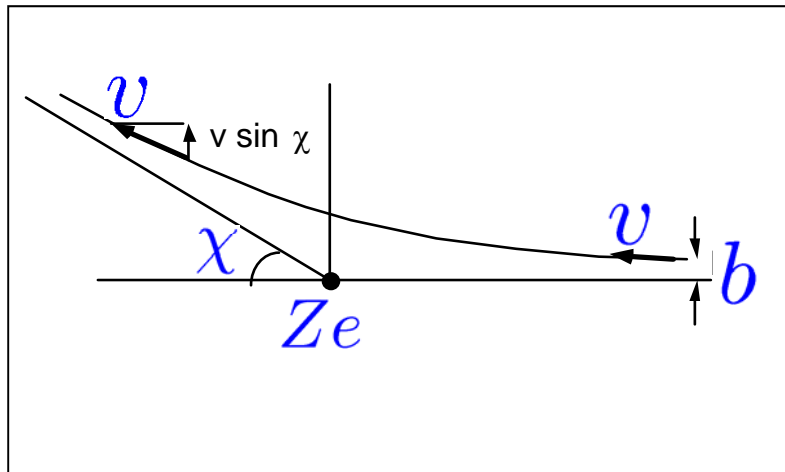
$$R = \underbrace{\langle \sigma v \rangle}_{\text{rate coefficient}} \times n$$

reaction rate (reactions per second)





# Coulomb collisions



cross section

$$\sigma = \pi b^2$$

(the area the colliding  
particle sees)

Coulomb cross section  
(pedestrian's approach)

$$E_{pot} = E_{kin}$$

$$\frac{1}{4\pi\epsilon_0} \frac{1}{b} = \frac{1}{2} m v^2 = \frac{3}{2} k_B T$$

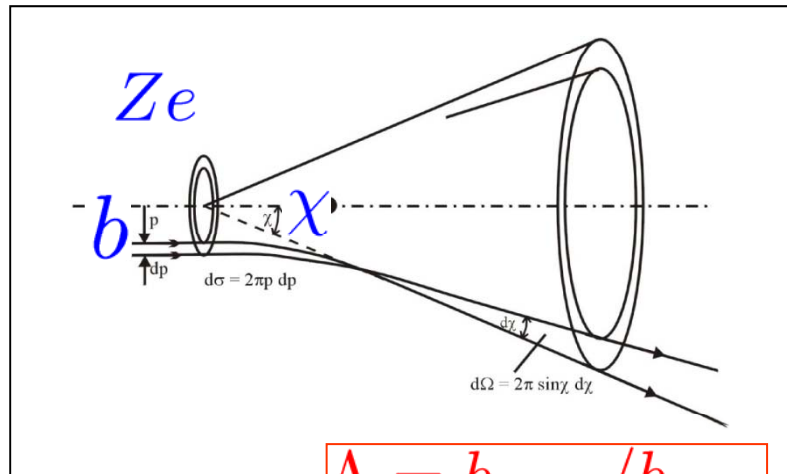
$$\sigma_{CB} = \frac{1}{36\pi\epsilon_0^2} \frac{e^4 Z^2}{k_B^2} \frac{1}{T^2}$$

$$\langle \sigma_{CB} v \rangle \sim T^{-3/2}$$

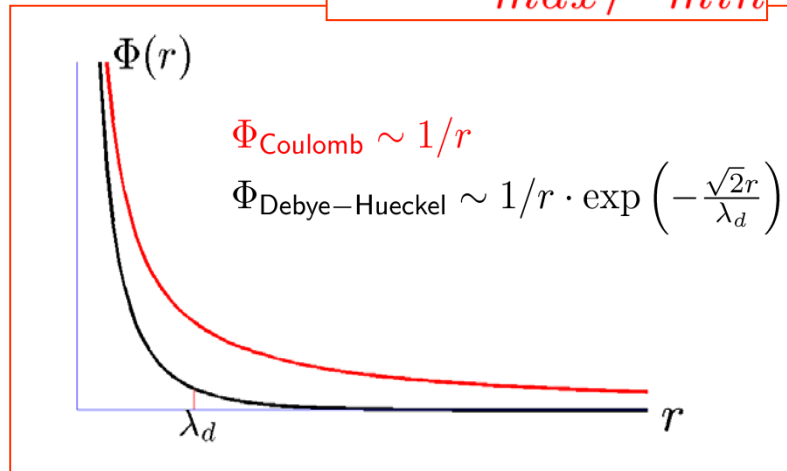
...the hotter, the less collisions!



# Coulomb collisions (contd.)



$$\Lambda = b_{max}/b_{min}$$



## Coulomb cross section (scattering theory)

+ integrate over all  $b$  (resp.  $\chi$ )

**but** Coulomb:  $1/r$  singularity

+ plasma: shielding  $\rightarrow b_{max} \approx \lambda_d$   
thermal motion  $\rightarrow b_{min} \approx \lambda_{\text{deBrog.}}$

$$\langle \sigma_{CB} v \rangle \sim T^{-3/2} \times \ln(\Lambda)$$

$\ln(\Lambda) \approx 15 \dots 20$  for fusion plasmas

small angle collisions are most effective  
( $\times 100$  for fusion plasmas)!



# Coulomb collisions (contnd.)



mean free path  $\lambda = \frac{1}{n\sigma}$   
 collision time  $\tau = \frac{\lambda}{v_{th}}$

$$\tau_{ee} \approx \tau_{ei} \approx \sqrt{\frac{m_e}{m_i}} \tau_{ii} \approx \frac{m_e}{m_t} \tau_{ie}$$

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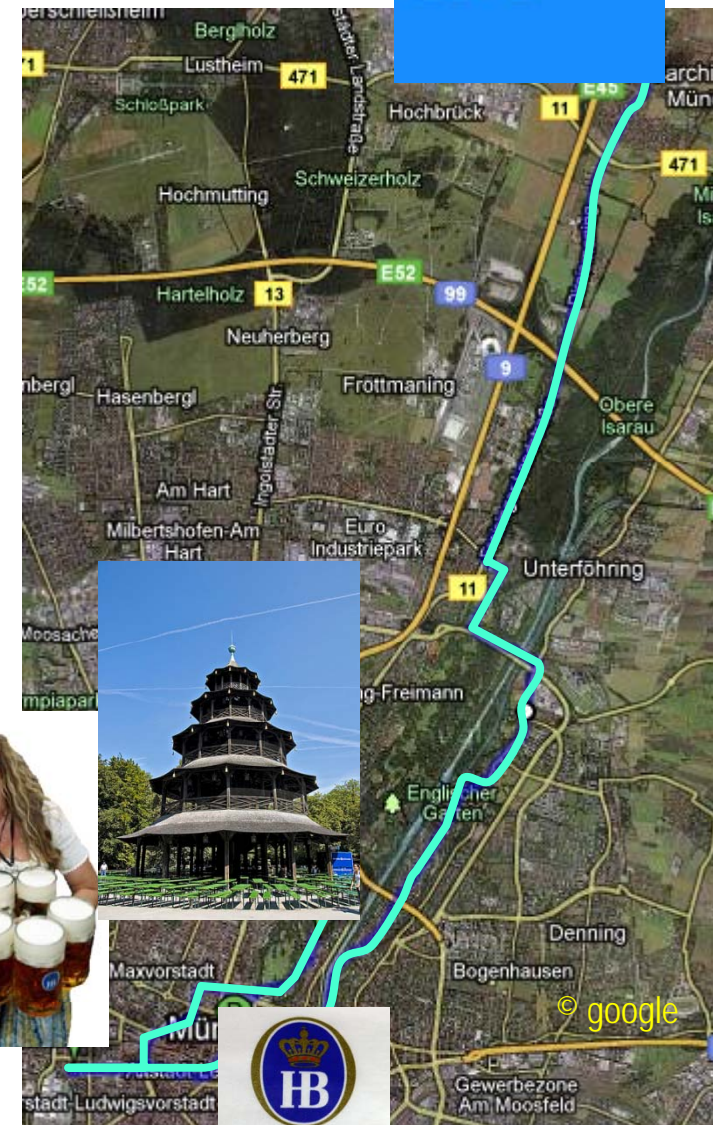
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$$T = 10 \text{ keV}, n = 1 \times 10^{20} \text{ m}^{-3}$$

collision	$\tau$	$\lambda_{CB}$
$e^- - T^+$	0.2 ms	9 km
$e^- - e^-$	0.3 ms	12 km
$D^+ - D^+$	18 ms	18 km
$T^+ - T^+$	23 ms	22 km

---

---







## Coulomb collisions (contd.)



resistivity

$$E = \eta j$$

$$\eta \approx \frac{m_e}{n e^2} \frac{1}{\tau_{ei}}$$

$$\eta(\Omega\text{m}) \approx 5.2 \times 10^{-5} Z \ln \Lambda (T(\text{eV}))^{-3/2}$$

figures:

$$\eta(T = 1.3\text{keV}) \approx 18 \text{ n}\Omega\text{m} \approx \eta_{\text{Copper}}$$

$$\eta(T = 10\text{keV}) \approx 0.8 \text{ n}\Omega\text{m} \approx 1/20 \times \eta_{\text{Copper}}$$

hot plasmas have a very low ohmic resistance!



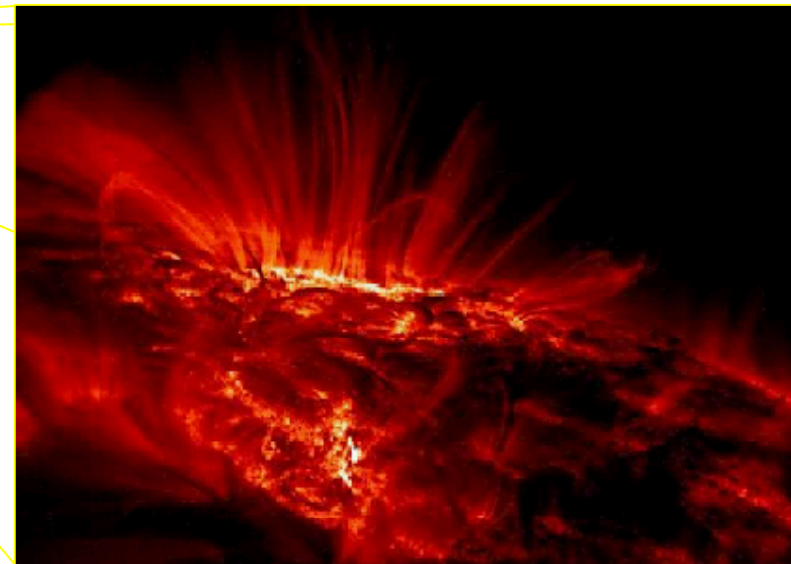
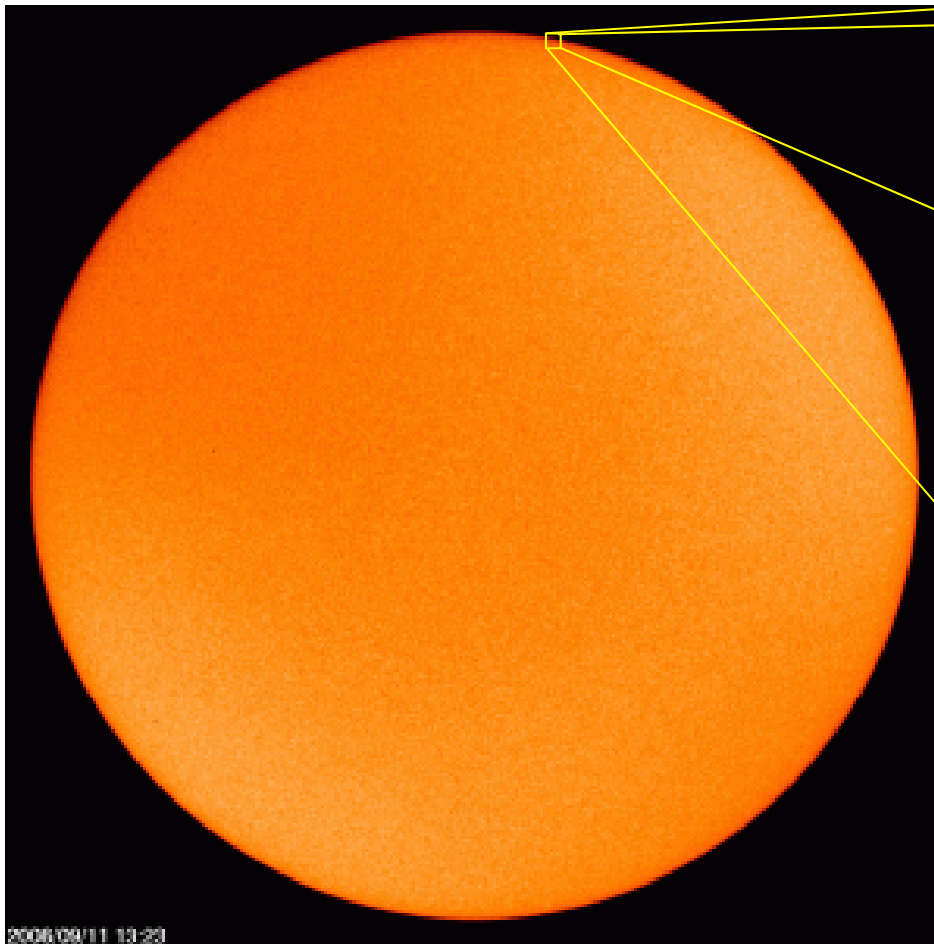
# Outline



- ① What is  $\pi\lambda\acute{\alpha}\sigma\mu\alpha$ ? The (A) world of plasmas.
- ② Overview: what makes a plasma a plasma?
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gravitational confinement

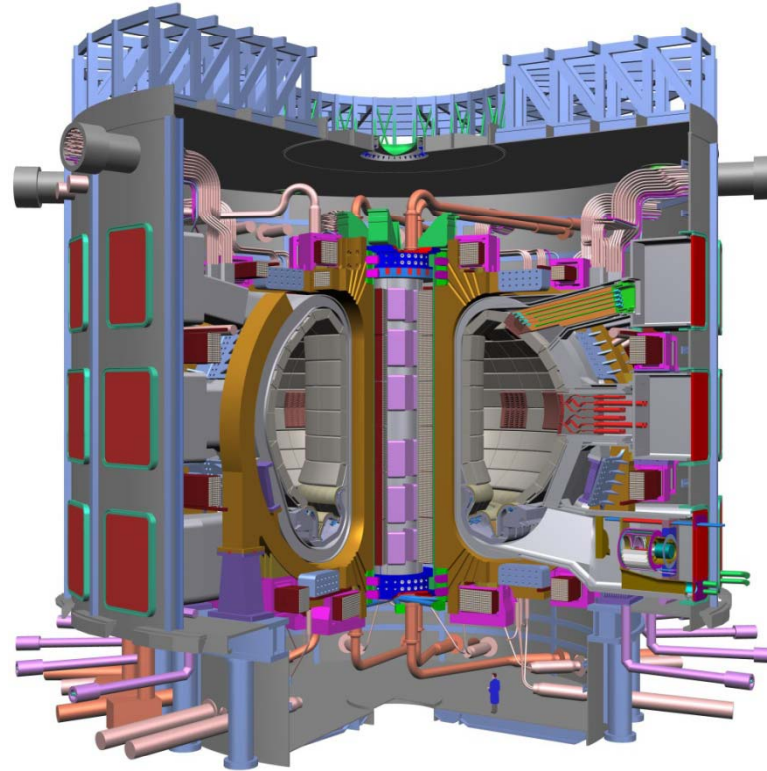


& electromagnetic forces

$$\vec{F} = m \frac{d\vec{v}}{dt} = q \left( \vec{E} + \vec{v} \times \vec{B} \right)$$



## How to get to a manageable size?



force plasma particles on toroidal orbits



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Motion in electro-magnetic fields

ideal plasma ( $N_d \gg \gg 1$ ) vs. single particle motion?

$$\lambda_{CB} \gg \gg L??$$

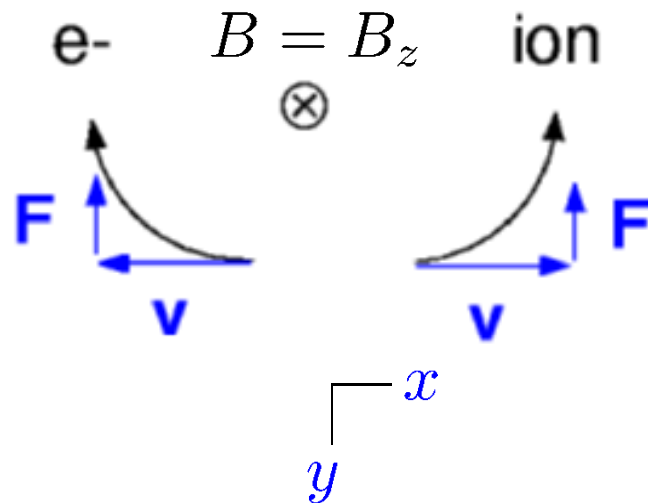


$$\mathbf{E} = 0, \mathbf{B} = \text{const}$$



- Lorentz force:

$$\vec{F} = m \frac{d\vec{v}}{dt} = q\vec{v} \times \vec{B}$$



$$m\dot{v}_x = qB_z v_y$$

$$m\dot{v}_y = -qB_z v_x$$

$$m\dot{v}_z = 0$$

$$\ddot{v}_{x,y} = -\left(\frac{qB}{m}\right)^2 v_{x,y}$$

$$v_z = \text{const}$$

- cyclotron frequency

$$\omega_c = \frac{qB}{m}$$

- Larmor radius

$$r_L = \frac{v_{\perp}}{\omega_c} = \frac{mv_{\perp}}{qB}$$



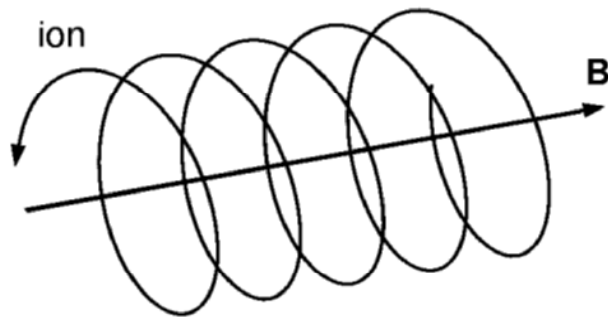


$$\mathbf{E} = 0, \mathbf{B} = \text{const}$$



- Lorentz force:

$$\vec{F} = m \frac{d\vec{v}}{dt} = q\vec{v} \times \vec{B}$$



helical trajectory along  $\mathbf{B}$

figures:

- cyclotron frequency for  $B = 2.5 \text{ T}$

$$f_{ce} = \omega_{ce}/(2\pi) \approx 70 \text{ GHz} \rightarrow \text{heating}$$

$$f_{cD} = \omega_{cD}/(2\pi) \approx 20 \text{ MHz}$$

- Larmor radius for  $T = 10 \text{ keV}$

$$\left. \begin{array}{l} r_{Le} \approx 95 \mu\text{m} \\ r_{LD} \approx 6 \text{ mm} \end{array} \right\} \ll L \ll \lambda_{CB}$$

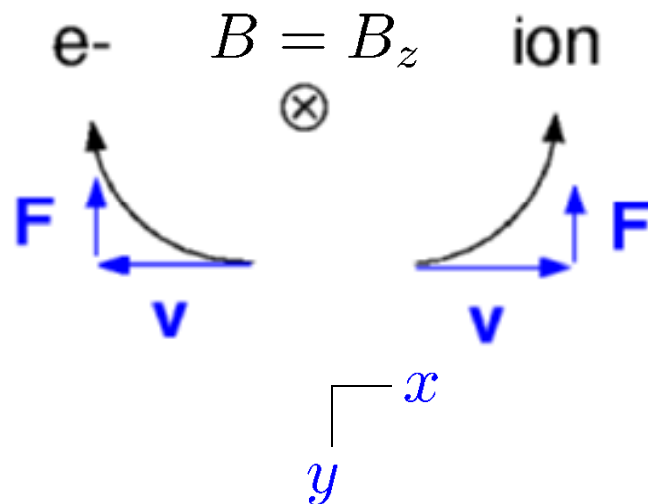


$$\mathbf{E} = 0, \mathbf{B} = \text{const}$$



- Lorentz force:

$$\vec{F} = m \frac{d\vec{v}}{dt} = q\vec{v} \times \vec{B}$$



Invariants:

- kinetic energy:

$$\frac{dW}{dt} = \frac{1}{2}m \frac{d}{dt}v^2 = m \vec{v} \cdot \frac{d\vec{v}}{dt}$$

$$= q \vec{v} \cdot (\vec{v} \times \vec{B})$$

$$\frac{dW}{dt} = 0$$

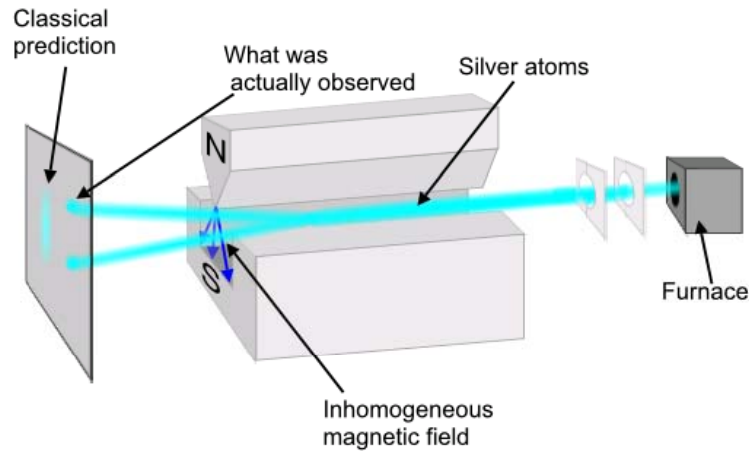
- $W$  is a **constant of motion** ( $\vec{B} = \text{const}$ )
- Lorentz forces guide, but do not work!
- in other words:  $\vec{F}_L$  pins particles to  $\vec{B}$
- inhomogeneous fields?



$$E = 0, B \neq \text{const}$$

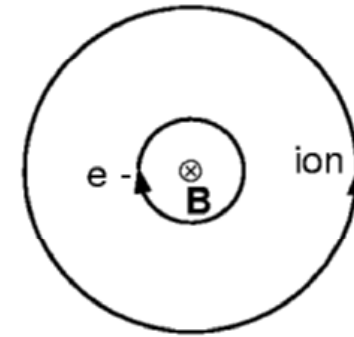


## Stern-Gerlach experiment



$$\vec{F} = -\mu \nabla_{\parallel} B$$

## gyro motion



magnetic moment  $\mu$

$$\mu = IA$$

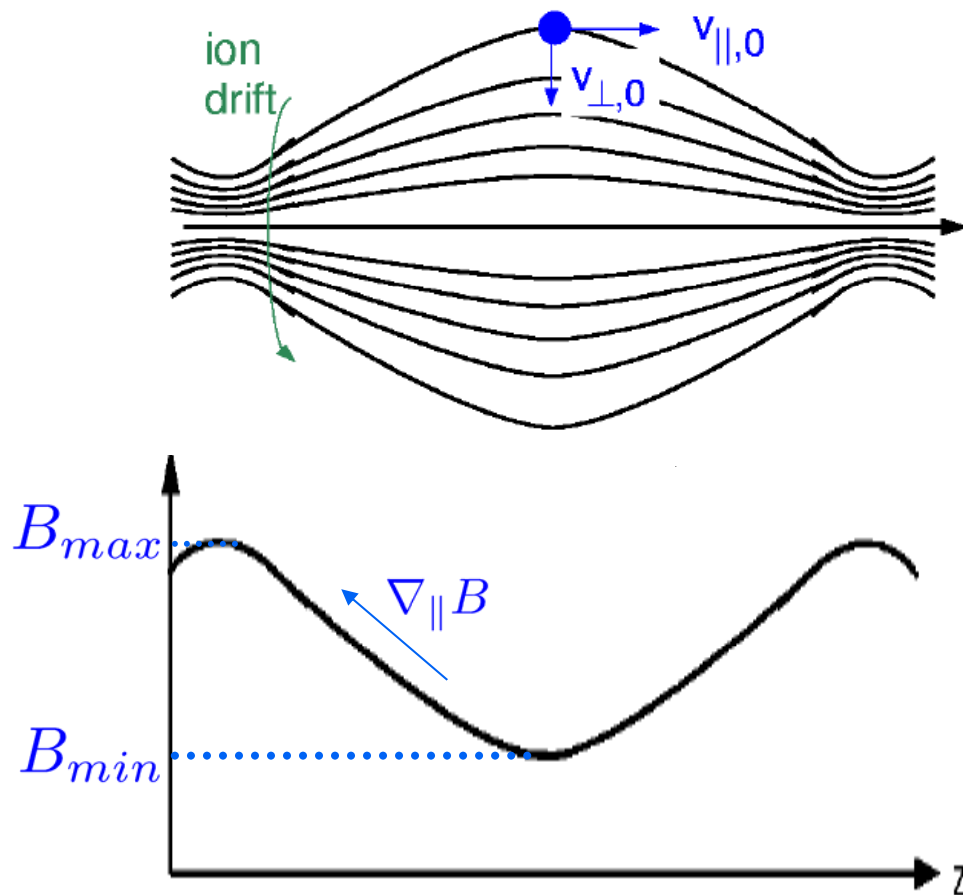
$$= \frac{e\omega_c}{2\pi} \pi r_L^2 = \frac{1}{2} \frac{mv_{\perp}^2}{B}$$



$E = 0$ ,  $B \neq \text{const}$  – App.



## Magnetic mirror



What's going on with  $\mu$ ?

$$m\dot{v}_{\parallel} = -\mu\partial_z B$$

$$m v_{\parallel} \dot{v}_{\parallel} = \frac{d}{dt} \left( \frac{1}{2} m v_{\parallel}^2 \right) = -\mu \partial_z B \frac{ds}{dt}$$

+ energy conservation  $\frac{d}{dt} \left( \frac{1}{2} m v_{\parallel}^2 + \mu B \right) = 0$

$$-\mu \frac{dB}{dt} + \frac{d}{dt} (\mu B) = 0$$

$$d\mu/dt = 0$$

$\mu$  is an *adiabatic invariant*

i.e. invariant for 'slow' ( $1/\omega_c$ ) proc.



# Adiabatic invariants



periodic motion, slow changes

$$\oint p dq = \text{const}$$

$p, q$ :

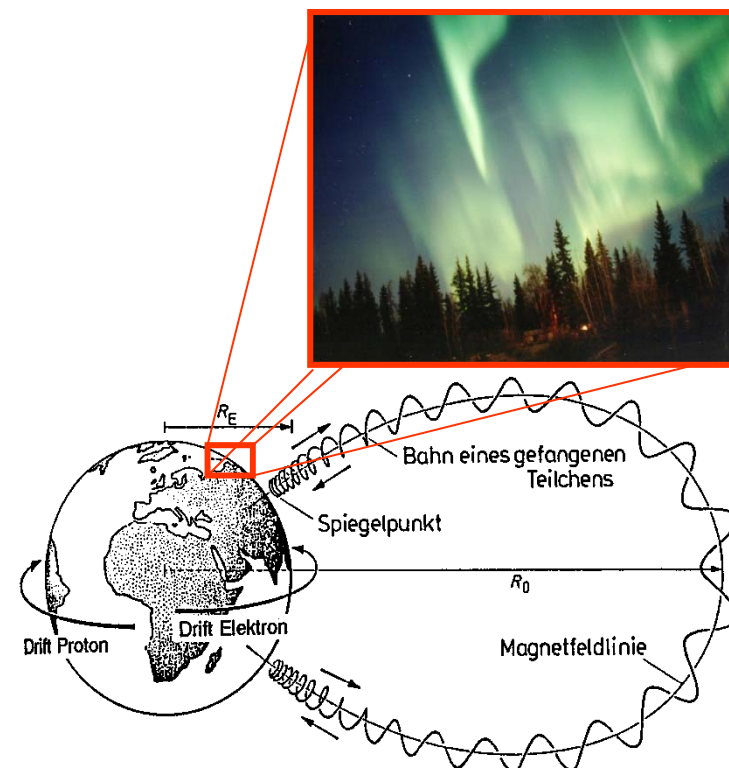
generalized momentum, coordinate

⇒ access to complex mechanics

applications in plasma theory:

- 1<sup>st</sup> inv.  $\mu$ : gyro-center motion
- 2<sup>nd</sup> inv.  $J$ : bounce motion →
- 3<sup>rd</sup> inv.  $\psi$ : precession

charged particles from **solar wind** are reflected in regions of stronger field strength and produce the polar light when exciting the atoms in the atmosphere





## Adiabatic invariants



periodic motion, slow changes

$$\oint p dq = \text{const}$$

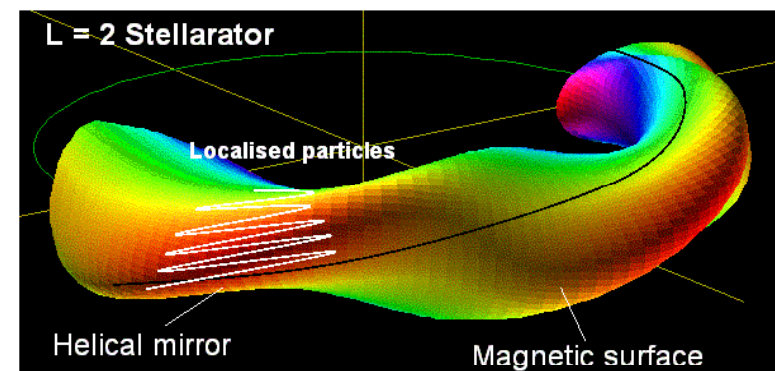
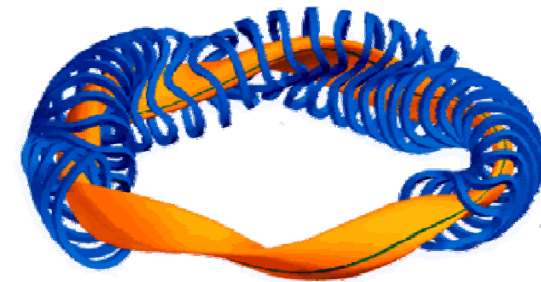
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- 2<sup>nd</sup> inv.  $J$ : bounce motion →
- 3<sup>rd</sup> inv.  $\psi$ : precession



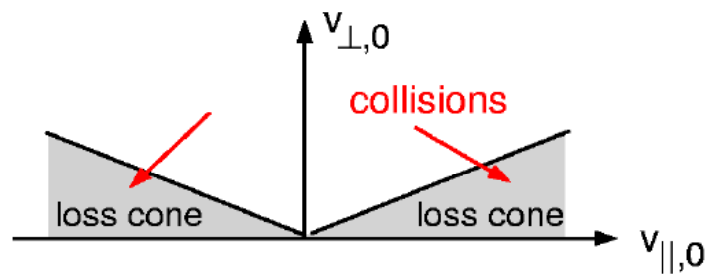
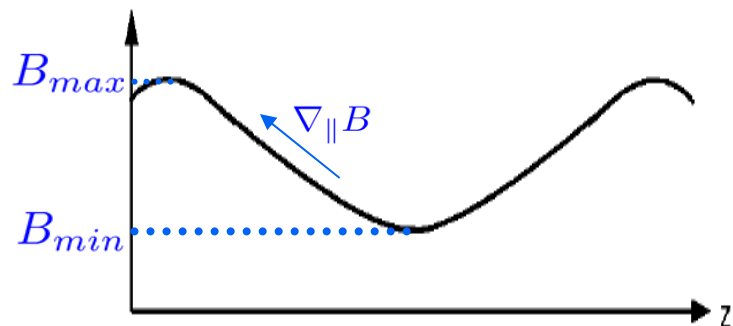




$E = 0, B \neq \text{const} - \text{App.}$



## Magnetic mirror



phase space

Confinement in magnetic mirrors?

$$\mu = \text{const} : \frac{1}{2} m v_{\perp \min}^2 / B_{\min} = \frac{1}{2} m v_{\perp r}^2 / B_{\text{refl.}}$$

$$W = \text{const} : v_{\perp r}^2 = v_{\perp \min}^2 + v_{\parallel \min}^2$$

$$\text{pitch angle: } \sin(\theta) = \frac{v_{\perp}}{v}$$

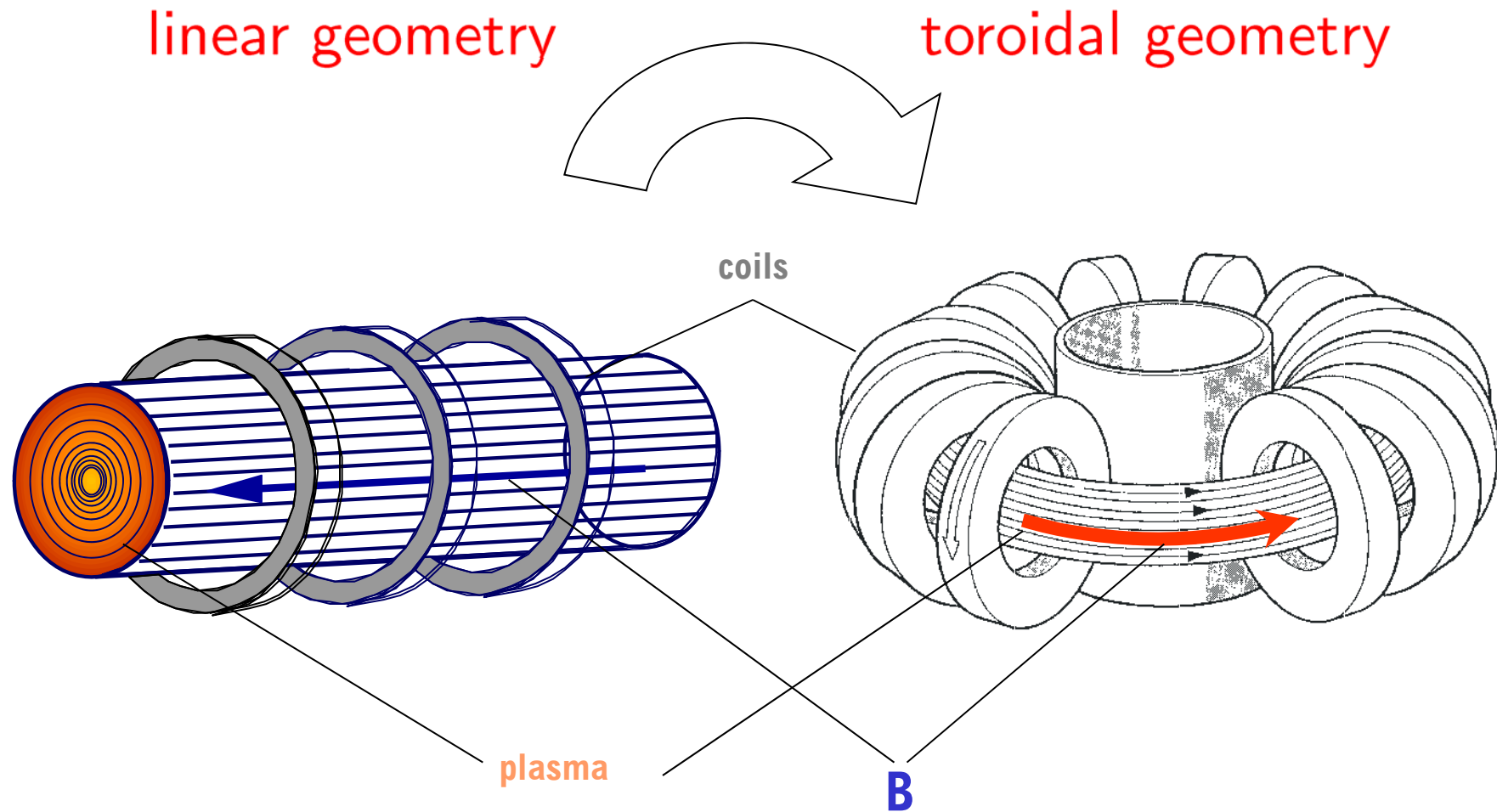
$$\sin^2(\theta_{\text{critical}}) = B_{\min} / B_{\max}$$

$\Rightarrow$  plasma lost at high  $v_{\parallel} / v_{\perp}$

No - how to avoid end losses?

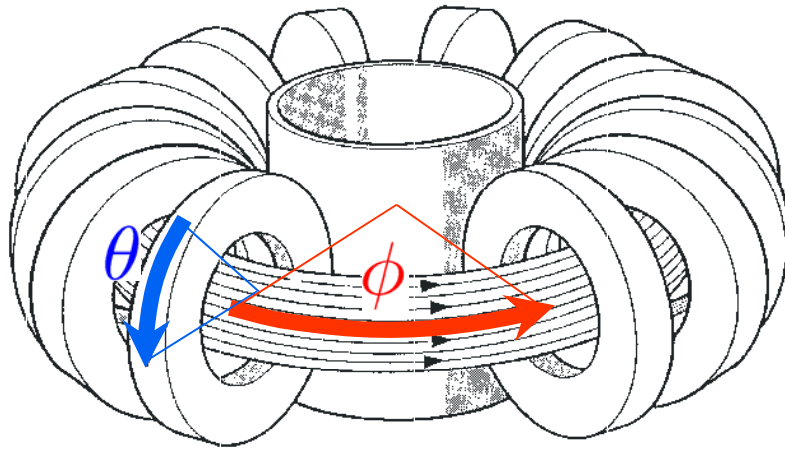


## overcoming losses at the ends





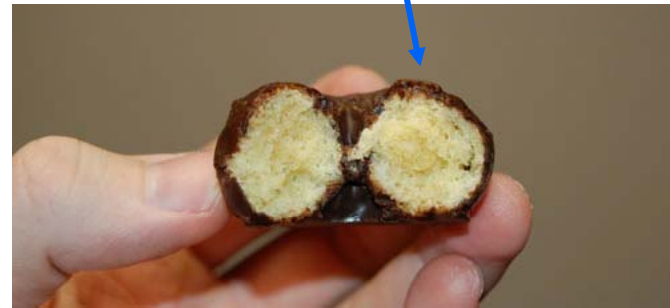
### torus and toroidal geometry



toroidal angle  $\phi$

poloidal angle  $\theta$

poloidal cross section ('cuts the pole')

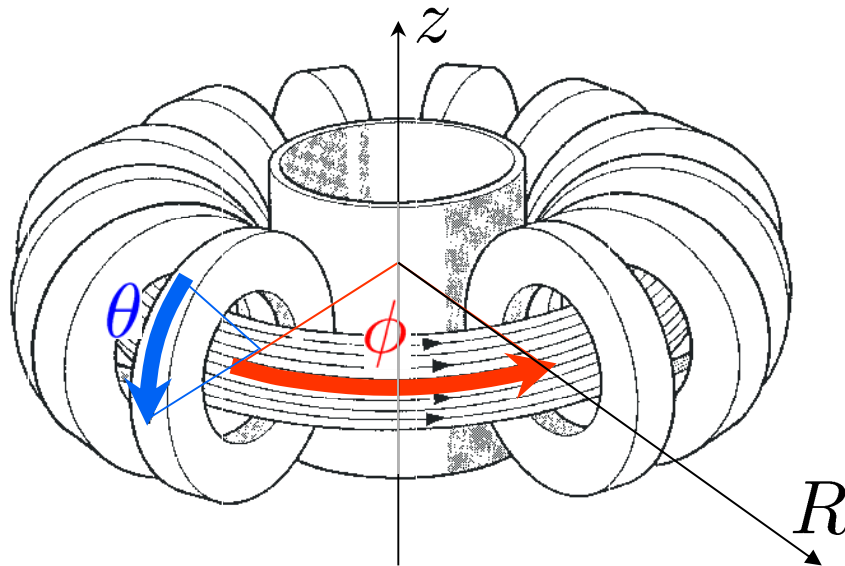


toroidal cross section





## field variation in toroids

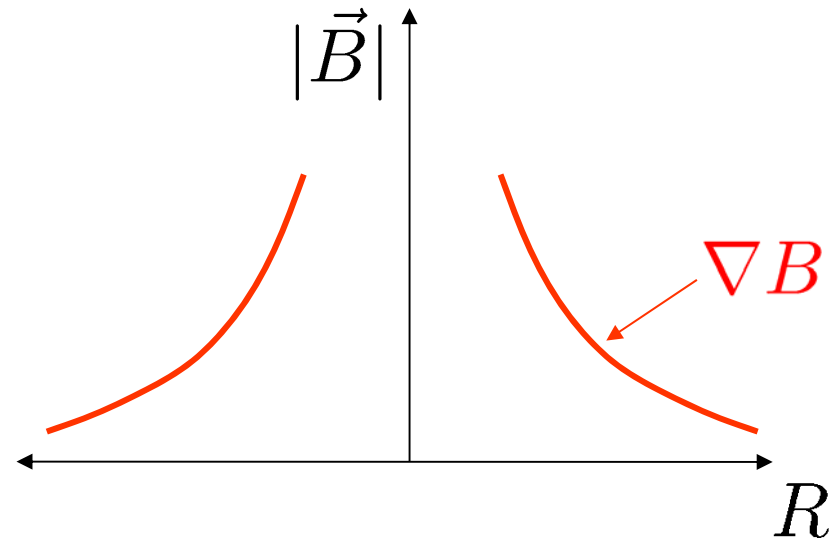
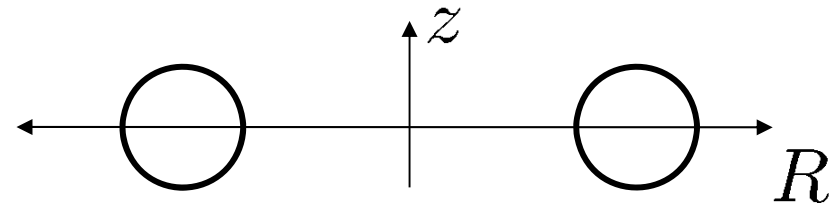


toroidal angle  $\phi$

poloidal angle  $\theta$

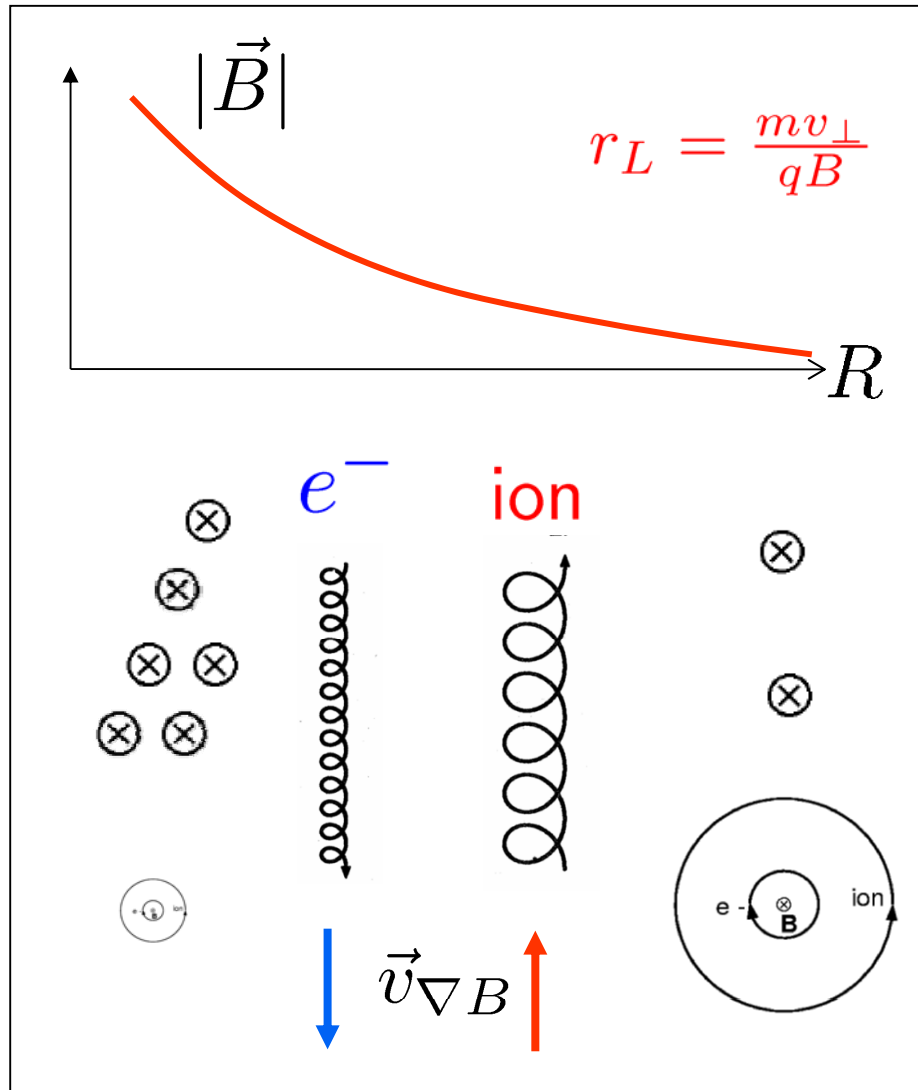
radius  $R$

poloidal cross section





## $\nabla B$ drift



force  $\rightarrow$  guiding centre drift

$$m\dot{\vec{v}} = \vec{F} + q\vec{v} \times \vec{B}$$

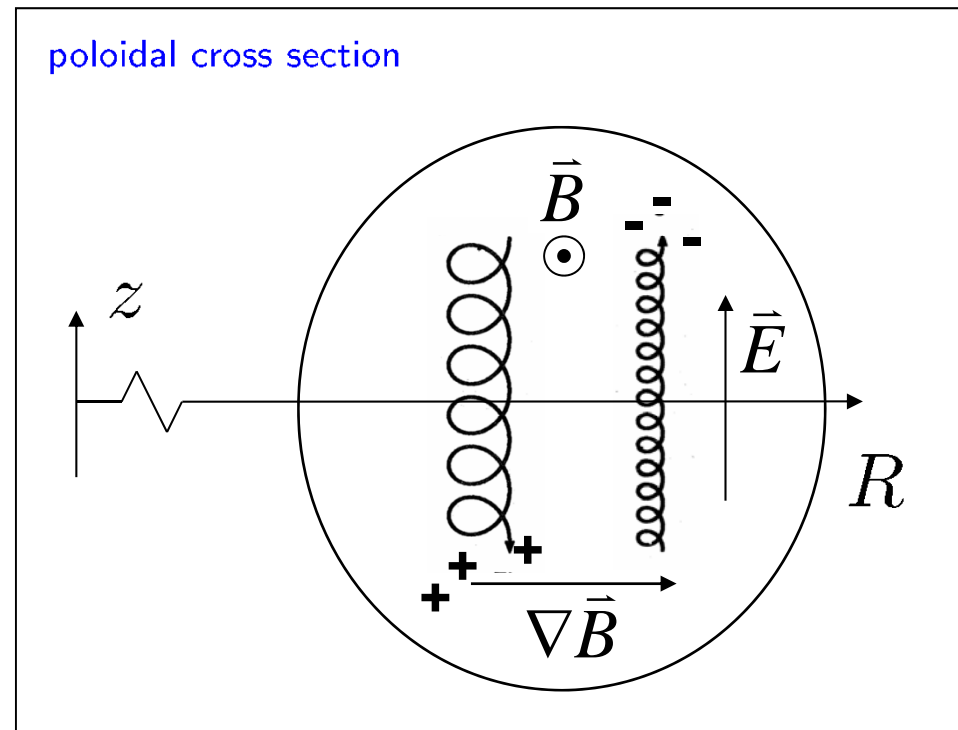
$$\vec{v}_D = \frac{1}{q} \frac{\vec{F} \times \vec{B}}{B^2}$$

here:  $\vec{F} = -\mu \nabla B$

$$\vec{v}_{\nabla B} = \pm \frac{1}{2} v_{\perp} r_L \frac{\nabla B \times \vec{B}}{B^2}$$



## $\nabla B$ drift in toroids



$\nabla B$ -drift polarizes the plasma toroid.

$$\Rightarrow \vec{E} = E \vec{e}_z$$

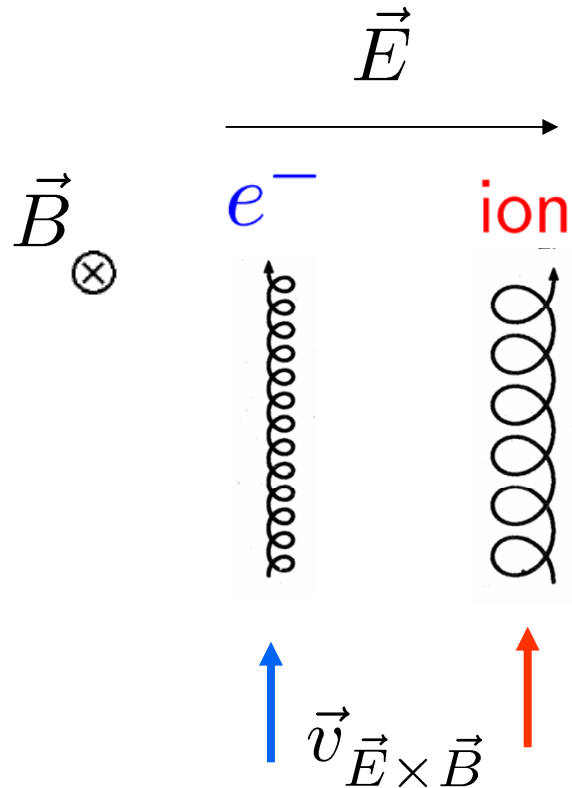




## $\vec{E} \times \vec{B}$ drift



$$r_L = \frac{mv_{\perp}}{qB}$$



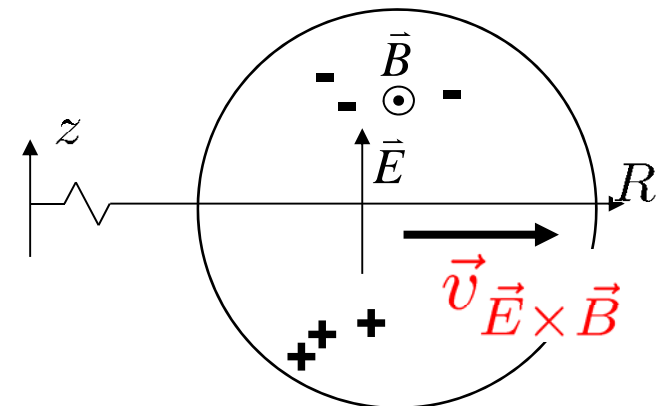
$$\vec{v}_D = \frac{1}{q} \frac{\vec{F} \times \vec{B}}{B^2}$$

force depends on  $q$

$$\vec{F} = q\vec{E}$$

$$\vec{v}_{\vec{E} \times \vec{B}} = \frac{\vec{E} \times \vec{B}}{B^2}$$

$\Rightarrow$  radial  $\vec{E} \times \vec{B}$ -drift in toroids



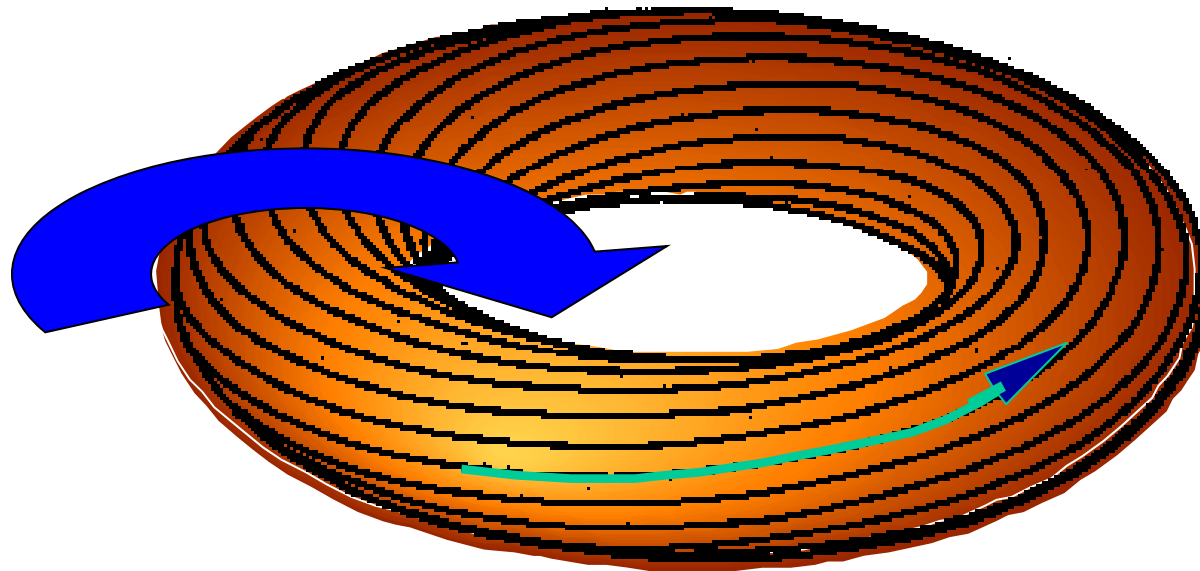


## Compensating drifts in toroids: rotational transform



compensation for drifts: twist of  $\vec{B}$

$$\vec{B} = \vec{B}_\phi + \vec{B}_\theta$$



$\vec{B}_\theta$

toroidal current  $\rightarrow$  tokamak  
external coils  $\rightarrow$  stellarator



## Compensating drifts in toroids: rotational transform



© M. Otte



## Summary (1/2)



A plasma is ...

- made of charged particles
- quasi-neutral many particles system
- roughly characterized by
  - plasma frequency  $\omega_p \sim \sqrt{n}$
  - Debye length  $\lambda_d \sim \sqrt{T/n}$



## Summary (2/2)



A fusion plasma ...

- must be hot ( $T_{DT} \approx 15 \text{ keV}$ )
- undergoes very rarely collisions
- can be confined by magnetic fields

Single-particle motion in magnetic fields:

- gyro-motion and (adiabatic) invariants
- drifts
- motion in toroidal fields

