



Aalto University
School of Science

Lecture 7: Collisions and Transport

Today's menu: weakly ionized gases

- Mean-free-path and collision frequency
- Mobility and *diffusion*
- Fick's law
- Sources & sinks: ionization & recombination
- Ambipolarity
- Decay times and steady-states
- Random walk and *diffusion*

Leaking out ...

- In real world, every vessel leaks
- So far we have assumed perfect confinement and infinite plasma
- In reality, plasma is finite → it has to have gradients
- Nature does not like gradients
 - *diffusion* from high to low density

What drives diffusion?

Collisions

Collisions in weakly ionized plasma

Weakly ionized plasmas – but why?

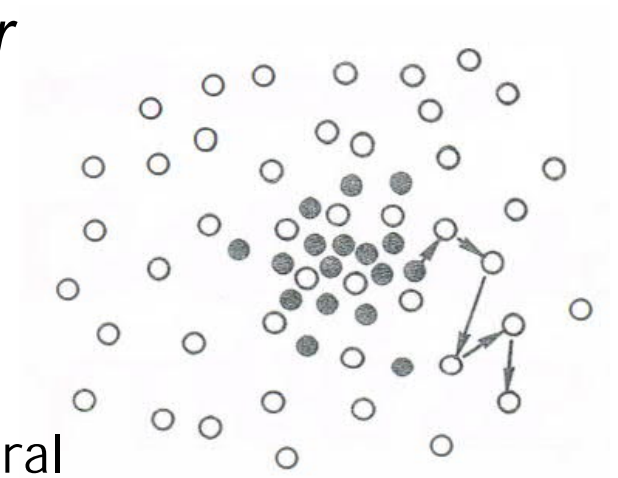
In fully ionized plasmas, collisions are *non-linear* effects

→ Mathematically complicated

→ Let's start with an easier case:

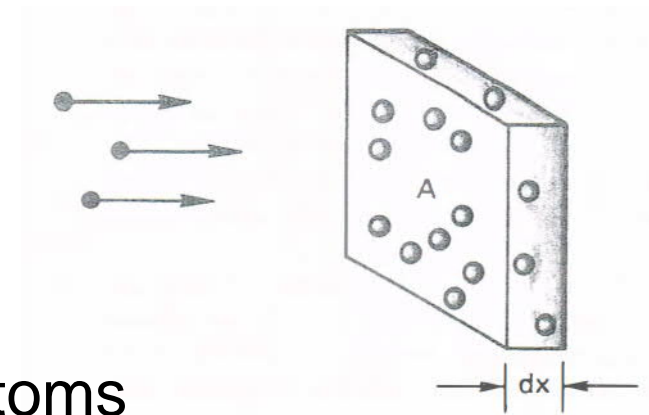
→ Study collisions in weakly ionized plasma

→ Charged particles suffer head-on collisions with neutral particles



Example: ionospheric plasma, $\frac{n_e}{n_n} \sim 10^{-6} - 10^{-3}$

Effect of collisions on flux



Flux Γ passes through a dense gas $\rightarrow \Gamma'$

- Dense gas consists of *scattering centers* = atoms
- Probability of colliding (= scattering of the flux) given by the *cross section* σ , which is the 'effective size' of an atom
- # of scatterers in a slab: $N = n_n \cdot A \cdot dx$
- Scatterers cover the fractional area $\frac{A_s}{A} = \frac{N \cdot \sigma}{A} = n_n \sigma dx$

$$\rightarrow \Gamma' = \Gamma - \Gamma \cdot \frac{N\sigma}{A} = \Gamma(1 - n_n \sigma dx)$$

'Freedom' parameters for plasma particles

$$\Gamma' - \Gamma = -\Gamma n_n \sigma dx \rightarrow \frac{d\Gamma}{dx} = -n_n \sigma \Gamma$$
$$\Gamma(x) = \Gamma_0 e^{-n_n \sigma x} \equiv \Gamma_0 e^{-x/\lambda_{mfp}}$$

Here, $\lambda_{mfp} \equiv 1/n_n \sigma$ is called the *mean-free path* for collisions

A related quantity is the *mean time* between collisions: $\tau = \frac{\lambda_{mfp}}{v}$

But:

- plasma particles have a distribution of velocities $\rightarrow \langle \sigma v \rangle = \int v \sigma(v) f(v) d^3v$
- Typically $\sigma = \sigma(v)$

$$\rightarrow \text{collision frequency: } \nu_{coll} = \frac{1}{\tau} = \frac{v}{\lambda_{mfp}} = n_n \langle \sigma v \rangle$$

Plasma motion due to collisions

Collisions cause friction → have to be included in the EoM:

$$mn \left[\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} \right] = qn\mathbf{E} - \nabla p - mn\nu_{coll}\mathbf{v}$$

Want to study effect of collisions *only* → simplify other stuff away:

1. Steady state
2. Low flow = assume friction dominates
3. Isothermal, $T = const$

$$\rightarrow \mathbf{v} = (qn\mathbf{E} - T\nabla n) / mn\nu_{coll} = \frac{q}{m\nu_{coll}} \mathbf{E} - \frac{T}{m\nu_{coll}} \frac{\nabla n}{n}$$

Diffusion in weakly ionized plasma

Our first transport coefficients ...

→ In the presence of collisions with *neutrals*, our plasma fluid moves according to the *density gradient* and *electric field*:

$$\Gamma_j = n\mathbf{v}_j = \pm\mu_j n\mathbf{E} - D_j\nabla n$$

where

$\mu_j \equiv \frac{q_j}{m_j\nu_{coll}}$ is called the *mobility* of the plasma

$D_j \equiv \frac{T_j}{m_j\nu_{coll}}$ is the *diffusion coefficient*

Important observations:

1. The flux is thus driven by *gradients*, as initially assumed:

$$\Gamma_j = \bar{\Gamma} \mu_j n \nabla \phi - D_j \nabla n$$

2. Collisions result into *diffusion* and diffusion in the presence of collisions means *transport*

Fick's law

For diffusion in regular gases the *Fick's law* applies

$$\Gamma = -D\nabla n$$

The physics of Fick's law:

- *Nature likes to flatten out gradients*
or, to put it in another way,
- *Gradients drive fluxes.*

A weakly ionized plasma thus obeys Fick's law ($E = 0$):

$$\Gamma_j = -D_j \nabla n$$

What is the time scale of flattening?

Fluids obey continuity equation:

$$\frac{\partial n_j}{\partial t} + \nabla \cdot \Gamma_j = 0$$

Quasineutrality $\rightarrow n_i \approx n_e \approx n \rightarrow \nabla \cdot \Gamma_e \approx \nabla \cdot \Gamma_i$

How about the individual fluxes?

Assume $\Gamma_e \neq \Gamma_i$

\rightarrow charge imbalance

\rightarrow electric field sufficient to retard electrons & accelerate ions to make $\Gamma_e = \Gamma_i$.

Ambipolar stuff ...

Find the magnitude of this *ambipolar electric field*:

$$\Gamma_e = \Gamma_i \rightarrow \mu_i n \mathbf{E} - D_i \nabla n = -\mu_e n \mathbf{E} - D_e \nabla n$$

$$\rightarrow \mathbf{E} = \frac{D_i - D_e}{\mu_i + \mu_e} \frac{\nabla n}{n}$$

\rightarrow The flux of the *plasma* is given by

$$\Gamma = \Gamma_i = \mu_i \frac{D_i - D_e}{\mu_i + \mu_e} \nabla n - D_i \nabla n = -\frac{\mu_e D_i + \mu_i D_e}{\mu_i + \mu_e} \nabla n \quad ; \text{Fick's law again!}$$

We have *ambipolar* fluxes driven by *ambipolar diffusion coefficient*

$$D_a \equiv \frac{\mu_e D_i - \mu_i D_e}{\mu_i + \mu_e} \approx D_i + \frac{\mu_i}{\mu_e} D_e \approx D_i + \frac{T_e}{T_i} D_i = D_i \left(1 + \frac{T_e}{T_i}\right)$$

Decay time of *weakly ionized plasma*

Now we have continuity equation for *the plasma*:

$$\frac{\partial n}{\partial t} + \nabla \cdot \mathbf{\Gamma} = \frac{\partial n}{\partial t} - D_a \nabla^2 n = 0$$

Recall Schrödinger eqn \rightarrow *separation of variables*: $n(\mathbf{r}, t) = X(\mathbf{r})T(t)$

Let's try to solve this in two simple geometries:

1. 1D case, i.e., slab geometry
2. 2D case, i.e., cylindrical geometry

Plasma decay time in slab geometry

Substitute trial fct to 1D continuity equation: $X(x) \frac{dT}{dt} - D_a T \frac{d^2 X}{dx^2} = 0$

$$\rightarrow \frac{1}{T} \frac{dT}{dt} = \frac{D_a}{X} \frac{d^2 X}{dx^2} = \text{const} \equiv -\frac{1}{\tau}$$

$$\rightarrow T(t) = n_0 e^{-t/\tau}$$

$$\rightarrow \frac{d^2 X}{dx^2} + \frac{1}{D_a \tau} X = 0 \rightarrow X(x) = A \sin kx + B \cos kx, \text{ where } k^2 \equiv \frac{1}{D_a \tau}$$

Plasma is bounded. Let boundaries be at $x = \pm L \rightarrow k = l\pi/2L$

$$\rightarrow n(x, t) = n_0 e^{-t/\tau} \cos \frac{\pi x}{2L}, \text{ why only } l = 1???$$

$$\rightarrow \text{the decay time is given by the diffusion coefficient: } \tau = \left(\frac{2L}{\pi}\right)^2 \frac{1}{D_a}$$

Sanity checks ...

Observations on τ :

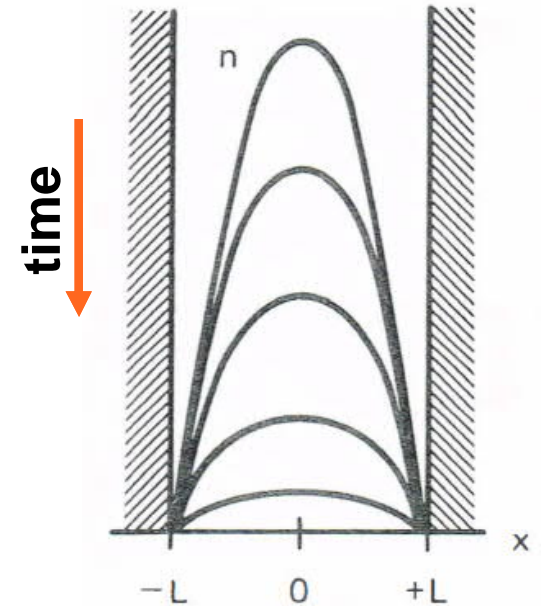
- τ increases with the box size L
- τ decreases with increasing diffusion

Makes sense. 😊

Also the shape of the solution, the *lowest diffusion mode*, looks reasonable, peaking at the center.



Weakly ionized plasma decays exponentially at rate determined by its size and the diffusion coefficient



The decay process

Start with an arbitrary initial shape

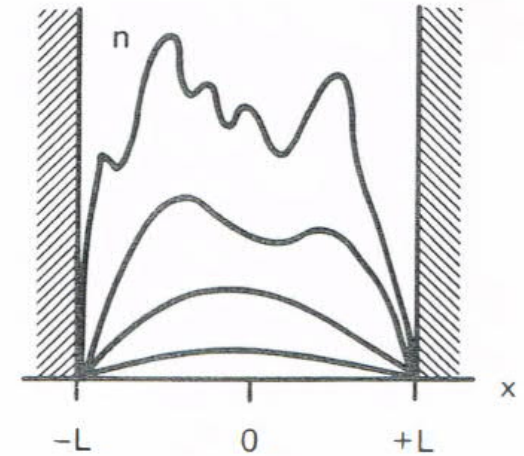
$$\text{FT} \rightarrow n(x, 0) = n_0 \left[a_0 + \sum a_l \cos \frac{(l + \frac{1}{2})\pi x}{L} + \sum b_m \sin \frac{m\pi x}{L} \right]$$

→ Trial solution:

$$n(x, t) = n_0 \left[a_0 e^{-t/\tau_0} + \sum a_l \cos \frac{(l + \frac{1}{2})\pi x}{L} e^{-t/\tau_l} + \sum b_m \sin \frac{m\pi x}{L} e^{-t/\tau_m} \right]$$

Substitute to the diffusion equation → $1/\tau_l = D_a \left[\left(l + \frac{1}{2} \right) \pi / L \right]^2$

→ $\tau_l = \left[\left(l + \frac{1}{2} \right) \pi / L \right]^{-2} 1/D_a$ → finest structures decay fastest!



Getting more realistic: Decay of a cylindrical plasma

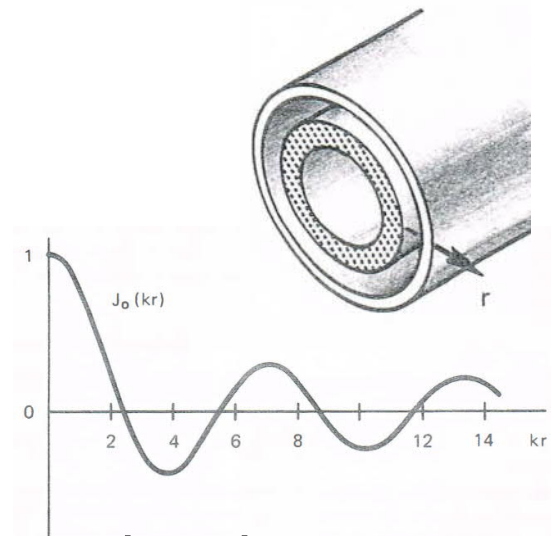
Assume cylindrical symmetry $\rightarrow \nabla^2 = \frac{\partial^2}{\partial r^2} + \frac{1}{r} \frac{\partial}{\partial r}$

Separate variables $\rightarrow \frac{d^2 X}{dr^2} + \frac{1}{r} \frac{dX}{dr} + \frac{1}{D\tau} X = 0$

In cylindrical geometry, the volume increase in r makes density drop faster \rightarrow could expect something like decaying cosine

Indeed, solutions are *Bessel functions!* Here, $J_0(r)$!

B.C's at $r = 0, r = a \rightarrow \frac{a}{\sqrt{D_a \tau}} = 2.4$ (first zero of J_0) $\rightarrow \tau = \left(\frac{a}{2.4}\right)^2 \frac{1}{D_a}$



How to get steady-state plasma...

... if plasma unavoidably decays due to inter-particle interactions?

Need a particle source, $S_+(\mathbf{r})$!!

Ways to 'feed' a plasma:

- Injection of particles
- Puffing of particles
- (recycling of particles – more about this later)

Simple steady-state cases:

1. local sources

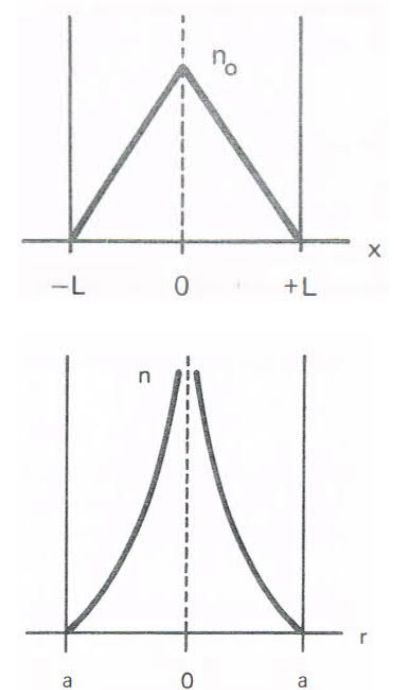
1-D case: a plane source at $x = 0$: $S_+(x) = S_+ \delta(0)$

→ For $x \neq 0$: $\frac{\partial^2 n}{\partial x^2} = 0 \rightarrow n(x) = n_0 \left(1 - \frac{|x|}{L}\right)$

2-D case: cylindrical plasma, line source at $r = 0$.

- (e.g., beam of energetic electrons causing ionization along the axis)

For $r \neq 0$: $\frac{1}{r} \frac{\partial}{\partial r} r \frac{\partial n}{\partial r} = 0 \rightarrow n(r) = n_0 \log \frac{a}{r}$, where $n(a) = 0$ was used



Simple steady-state cases: 2. ionization source

Plasma can be fuelled also by a *heat* source (in cold plasmas):
electrons in the hot Maxwellian tail keep ionizing the gas neutrals

a 'continuous' source (around heat source): $S_+ \propto n$.

Let's write then $S_+(\mathbf{r}) = Zn(\mathbf{r})$, where $Z \neq Z(\mathbf{r})$ is the *ionization* fct

$$\rightarrow \nabla^2 n = -\frac{Z}{D}n$$

But this is formally the same as the eqn for $X(r) \rightarrow n(r) = J_0(r)$

How about *sinks*?

We just had *ionization* as a source.

The reverse process, *recombination*, is a sink, S_- .

Recombination requires both electrons and ions $\rightarrow S_- \propto n_i n_e$.

Study the effect of recombination alone = neglect diffusion

$\rightarrow \frac{\partial n}{\partial t} = -\alpha n^2$, where α is the recombination coefficient, $\alpha \neq \alpha(n)$

Non-linear equation! \rightarrow separation of variables not possible

\rightarrow solution by 'eye-balling': $\frac{1}{n(r,t)} = \frac{1}{n_0(r)} + \alpha t$ (HW: just show)

New processes can change the character of the solutions

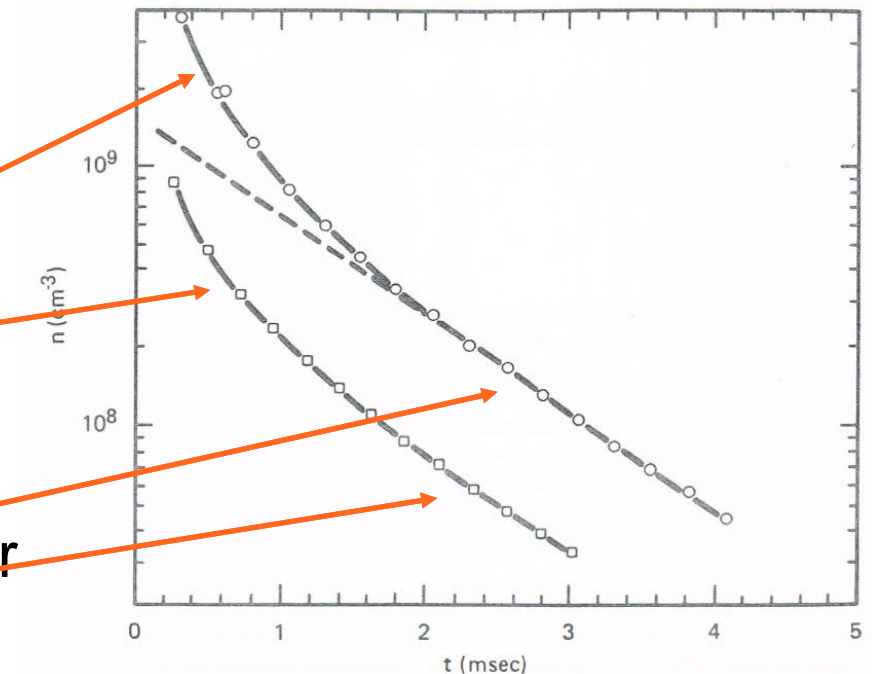
At high density, recombination ($\propto n^2$) typically dominates

$$\rightarrow n(r, t) \propto \frac{1}{\alpha t}$$

and the density falls *reciprocally* in time, *not* exponentially!

As the density drops, diffusion takes over

\rightarrow exponential decay



Until now, we have been studying 'freely floating' plasmas

But mostly we are interested in *magnetized* plasmas!

How does the plasma decay when it is imbedded in a confining magnetic field?

Like *fusion* or *atmospheric* or *solar* plasmas...

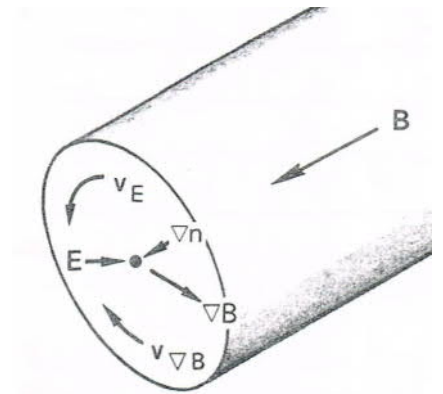
What does the magnetic field do in weakly ionized plasmas?

In direction parallel to \mathbf{B} , magnetic field has no say
→ same physics as before

What is interesting is the transport *perpendicular* to \mathbf{B} .
These particles are glued to the fieldlines.

... But we can have cross-field drifts! $\mathbf{E} \times \mathbf{B}$ & Co!

Luckily drifts can be aligned so that they are parallel to walls (laboratory plasmas)



Analyze fluid equations $\perp \mathbf{B} = B_0 \hat{\mathbf{z}}$

Same simplifying assumptions as before \rightarrow

$$\text{Motion } \perp \mathbf{B} : mn \frac{d\mathbf{v}_\perp}{dt} \approx 0 \approx nq(\mathbf{E} + \mathbf{v} \times \mathbf{B}) - T\nabla n - mn\nu_{coll}\mathbf{v}_\perp$$

$$v_x = \pm \mu E_x - \frac{D}{n} \frac{\partial n}{\partial x} \pm \frac{\Omega}{\nu_{coll}} v_y$$

$$v_y = \mp \mu E_y - \frac{D}{n} \frac{\partial n}{\partial y} \mp \frac{\Omega}{\nu_{coll}} v_x$$

\rightarrow HW:

$$\mathbf{v}_\perp = \pm \mu_\perp \mathbf{E} - D_\perp \frac{\nabla_\perp n}{n} + \frac{\mathbf{v}_{E \times B} + \mathbf{v}_{dia}}{1 + \nu_{coll}^2 / \Omega^2},$$

where $\mu_\perp \equiv \mu / (1 + \Omega^2 \tau_{coll}^2)$ and $D_\perp \equiv D / (1 + \Omega^2 \tau_{coll}^2)$

Physics of v_{\perp}

1. Familiar magnetic drifts perpendicular to their respective gradients ($\mathbf{v}_{E \times B} \propto \nabla \phi$, $\mathbf{v}_{dia} \propto \nabla n$), but slowed down by collisions with neutrals by the *drag factor* $1 + v_{coll}^2 / \Omega^2$.

- Increase magnetic field and/or reduce neutral density \rightarrow good old drifts!

2. Mobility drift parallel to \mathbf{E} and diffusion drift parallel to ∇n , obtained in the absence of \mathbf{B} are now slowed down by the factor

$$1 + \Omega^2 \tau_{coll}^2$$

- This is not the same as the drag factor but works the opposite way (as it should): increase magnetic field and/or reduce neutral density \rightarrow mobility and diffusion drifts vanish

More on physics of v_{\perp} -- random walk ...

$\Omega\tau_{coll} \ll 1 \rightarrow$ B-field has little effect on diffusion

$\Omega\tau_{coll} \gg 1 \rightarrow$ B-field reduces diffusion across B

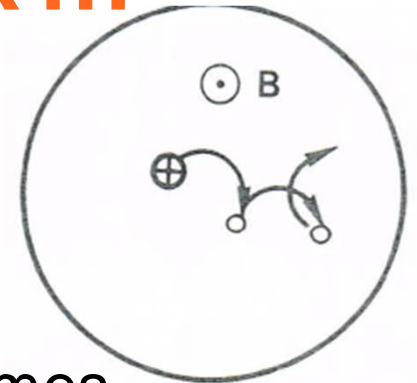
The physics of 'magnetic' slowing down of diffusion:

In the presence of strong B the diffusion coefficient becomes

$$D_{\perp} \rightarrow \frac{T}{m\nu_{coll}} \frac{1}{\Omega^2\tau_{coll}^2} = \frac{T}{m\Omega^2} \nu_{coll}$$

We then realize: $\frac{T}{m\Omega^2} \sim \frac{v_{th}^2}{\Omega^2} = r_L^2 \rightarrow D_{\perp} \sim r_L^2 \nu_{coll} \sim \text{stepsize}^2 / \text{colltime}$

\rightarrow the effect of B is to reduce the step size from mean-free-path to Larmor radius!



Differences to 'free-floating' plasma

No \mathbf{B} -field (or parallel to it): collisions *retard* the motion

$$\rightarrow D \propto 1/\nu_{coll}$$

Across the \mathbf{B} -field; collisions are *needed* for particles to jump from one Larmor orbit to another

$$\rightarrow D \propto \nu_{coll}$$

Also the role of particle mass is reversed:

- No \mathbf{B} (or $\parallel \mathbf{B}$): $D \propto 1/\sqrt{m}$; *light electrons move faster along \mathbf{B}*
- $\perp \mathbf{B}$: $D \propto \sqrt{m}$; *ions have larger Larmor radius = step size*